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## Thin Solid Films

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# Structural and optical properties of Cu<sub>2</sub>ZnSnS<sub>4</sub> thin film absorbers from ZnS and Cu<sub>3</sub>SnS<sub>4</sub> nanoparticle precursors

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#### ABSTRACT

 $Cu_2ZnSnS_4$  (CZTS) has been considered as an alternative absorber layer to  $Cu(In,Ga)Se_2$  due to its earth abundant and environmentally friendly constituents, optimal direct band gap of 1.4–1.6 eV and high absorption coefficient in the visible range. In this work, we propose a solution-based chemical route for the preparation of CZTS thin film absorbers by spin coating of the precursor inks composed of  $Cu_3SnS_4$  and ZnS NPs and annealing in  $Ar/H_2S$  atmosphere. X-ray diffraction and Raman spectroscopy were used to characterize the structural properties. The chemical composition was determined by energy dispersive X-ray spectroscopy. Optical properties of the CZTS thin film absorbers were studied by transmission, reflection and photoluminescence spectroscopy.

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#### 1. Introduction

Cu<sub>2</sub>ZnSnS<sub>4</sub> (CZTS) has been considered as an alternative absorber layer to Cu(In,Ga)Se2 due to its earth abundant and environmentally friendly constituents, optimal direct band gap of 1.4-1.6 eV and high absorption coefficient ( $>10^4$  cm<sup>-1</sup>) in the visible range [1,2]. In recent years, great efforts have been focusing on the preparation of CZTS thin films and exploration of their potential application in thin film solar cells [2-10]. The preparation of CZTS thin films can be summarized as two methods, namely vacuum-based processes and non-vacuum solution-based processes [11]. More specifically, vacuum-based processes include sputtering and evaporation while non-vacuum based processes include electrodeposition, spray pyrolysis, and ink-based approaches [2–10,12]. The solution processes such as ink rolling and printing which allow the large scale deposition of thin films are considered as one of the low cost routes for fabrication of electronic devices. Cu2ZnSnSe4 and Cu2ZnSn(S, Se)4-based solar cells using a hydrazine-based solution process have already reached an energy conversion efficiency as high as 10.1% [3,4], demonstrating the effectiveness of the solution processes in CZTSe-based solar cells. In this contribution, we propose a solution-based chemical route for the preparation of CZTS thin film absorbers from Cu<sub>3</sub>SnS<sub>4</sub> and ZnS nanoparticle precursors. The structural and optical properties have been studied by X-ray diffraction, Raman spectroscopy, transmission, reflection and photoluminescence spectroscopy.

#### 2. Experimental details

#### 2.1. Preparation of $Cu_3SnS_4$ and ZnS nanoparticle inks

ZnS NPs (NPs) were prepared according to the literature [13] with some modification. Briefly, stoichiometric zinc acetate dihydrate and sulfur were mixed together with oleylamine and heated to 240 °C to allow formation of ZnS NPs. Cu<sub>3</sub>SnS<sub>4</sub> NPs were synthesized by one pot techniques based on the method reported on reference [14] where copper acetylacetonate, tin chloride, sulfur and oleylamine were mixed together in one vessel and heated to 250 °C to allow the reaction to take place. In this approach, oleylamine was used as both solvent and stabilizer. The inks were formed by mixing certain amount of ZnS and Cu<sub>3</sub>SnS<sub>4</sub> NPs dispersed in hexanethiol.

#### 2.2. Preparation of thin films

Thin films were deposited on different substrates (Mo/glass, 10 nm Sn on Mo/glass and 20 nm Sn on Mo/glass substrates) by spin coating using the precursor inks, followed by sequential heat treatment at 200 °C for 5 min in air to remove the solvent. After that, the resulting precursor thin films were annealed at 540 °C under Ar/H<sub>2</sub>S (5%)-atmosphere for 60 min to allow the formation of CZTS absorber by the reaction of Cu<sub>3</sub>SnS<sub>4</sub> and ZnS NP precursors. In order to study the optical transmission and reflection properties, corresponding samples on soda lime glass and tin layer coated soda lime glass were prepared. Note that, prior to the film coating, the glass and Mo/glass substrates were cleaned by subsequent ultrasonication in acetone, ethanol and distilled water, each for 15 min and dried under nitrogen stream.

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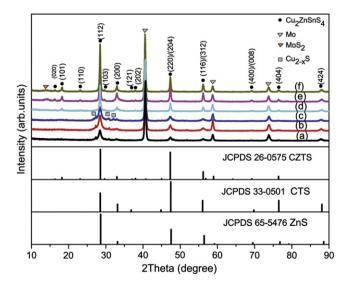
#### 2.3. Characterization

The structure of the films was studied by X-ray diffraction (XRD) and Raman spectroscopy. XRD were operated in the 2θ range from 10 to 90° on a Bruker D8-Advance X-ray diffractometer with CuK $\alpha$  radiation ( $\lambda = 1.5406 \text{ Å}$ ) using a step size of 0.02° and step time of 0.3 s. For the Raman measurement a Ti:Sapphire-ring-laser was used as an excitation. The wavelength of the laser is fully tenable from 690 nm to 1050 nm. To avoid laser heating the beam power was kept below 3.5 mW. Raman spectra were recorded with a Horiba T64000 triple monochromator system in backscattering configuration with a microscope and a motorized XY stage. The micro-Raman spectroscopy with a 100× objective was performed at room temperature with a wavelength of 747 nm. The chemical compositions were obtained by energy dispersive X-ray spectroscopy analysis (EDX) which was performed in a LEO GEMINI 1530 field-emission gun scanning electron microscope (SEM) with the operating voltage of 10 kV and a Thermo Noran X-ray silicon drift detector (acquisition and evaluation software Noran System Seven). The absorption was characterized through transmission and reflection measurements performed with a Lambda 950 UV-vis spectrometer. Photoluminescence (PL) was performed at room temperature using an excitation diode laser with a wavelength of 660 nm and a silicon CCD camera.

#### 3. Results and discussion

#### 3.1. Structural properties

Fig. 1 shows the XRD patterns of as-deposited and annealed samples on different substrates. Bragg peaks belonging to phases of  $Cu_{2-x}S$  can be observed in the XRD patterns of as-deposited samples, which may be due to the decomposition of  $Cu_3SnS_4$  precursors during the heat treatment process. However, these Bragg peaks disappear after annealing. The appearance of the Bragg peaks of (020), (101), (110), (103) and (202) which belong to kesterite CZTS (JCPDS data file no.: 26-0575(CZTS)) were observed in all three annealed samples, indicating the formation of kesterite phase of CZTS. Although two of the samples were prepared on molybdenum coated glass substrates with 10 and 20 nm tin layer on top we cannot find any Bragg peaks corresponding to  $Sn_xS_y$  after annealing at 540 °C. This suggests that the tin layer converted to CZTS by reaction with  $H_2S$ , ZnS and



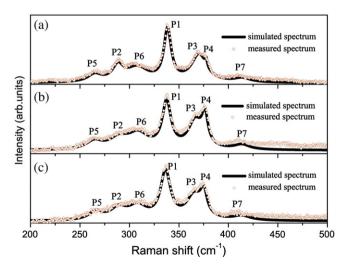
**Fig. 1.** X-ray diffraction patterns of: as-deposited precursor thin films prepared on (a) Mo/glass substrate, (b) Mo/glass substrate coated with 10 nm Sn layer, (c) Mo/glass substrate coated with 20 nm Sn layer and corresponding annealed samples: (d), (e), and (f). For reference, the XRD patterns of CZTS (JCPDS 26–0575),  $Cu_3SnS_4$  (JCPDS 33–0501) and ZnS (65–5476) are shown below.

Cu<sub>3</sub>SnS<sub>4</sub>. In addition to the kesterite CZTS and Mo Bragg peaks, peaks ascribed to MoS<sub>2</sub> phase were also identified. Note that the relative intensity of MoS<sub>2</sub> Bragg peak is much more intensive in Mo coated with tin layer samples than in the bare Mo substrate sample. This phenomenon can be explained as the better penetration of H<sub>2</sub>S through the tunnel created by the diffusion of tin underneath the nanoparticle layer during annealing process. It is known that the main XRD peaks among kesterite CZTS, ZnS and CTS are overlapping, as shown in Fig. 1. Therefore, although no noticeable peak ascribed to the other secondary phases can be observed except for the Mo and MoS<sub>2</sub> Bragg peaks we cannot conclude that both ZnS and CTS NPs precursors have reacted completely and transformed to CZTS.

In order to find out the existence of the secondary phases, Raman spectroscopy was performed. Fig. 2 shows the Raman spectra of annealed samples on different substrates. The strongest peak P1 in all three samples is located at 337–338 cm<sup>-1</sup>, which is identified as the main vibrational A symmetry mode from kesterite CZTS [9–11,15,16]. The weaker peaks at about 264–266, 288, and 364–375 cm<sup>-1</sup> were also found by fitting the spectra with Lorentzian curves, and can be attributed to kesterite CZTS [9-11,15,16]. Note that the peak at 364-375 cm<sup>-1</sup> is split into two peaks, P3 and P4 centered at 364–368 and 374–377 cm<sup>-1</sup>, respectively. This may be due to the use of excitation laser wavelength of 747 nm during the measurements, which is in agreement with the report by Fernandes et al. [10] who studied the effect of excitation laser wavelength on the Raman spectra of CZTS and found that the peak at 368 cm<sup>-1</sup> separated into two peaks when using an excitation laser with a wavelength of 785 nm. Moreover, the Raman spectra in all three samples also show additional weaker broad peaks P6 in the region of 305-315 cm<sup>-1</sup>. The possible contribution of the broad peaks could come from the phase of cubic Cu<sub>2</sub>SnS<sub>3</sub>, SnS<sub>2</sub> and Cu<sub>3</sub>SnS<sub>4</sub> which are reported to have vibration modes at 303, 314 and 318 cm<sup>-</sup> respectively [17,18]. However, the presence of SnS<sub>2</sub> can be ruled out due to the absence of the Bragg peaks of SnS2 in XRD patterns. The peak P7 located at 411-413 cm<sup>-1</sup> indicates the presence of MoS<sub>2</sub>, which further confirmed the XRD results [15].

#### 3.2. Chemical composition

The chemical compositions of the as-deposited and annealed thin films on Mo substrates were studied using EDX measurements, as shown in Table 1. Due to the overlap between sulfur  $K\alpha$  peak and molybdenum  $L\alpha$  peak, EDX cannot resolve sulfur from molybdenum.



**Fig. 2.** Raman spectra of annealed CZTS thin films prepared on different substrates: (a) Mo/glass substrate, (b)Mo/glass substrate coated with 10 nm Sn layer, and (c) Mo/glass substrate coated with 20 nm Sn layer. Seven peaks were observed in all three samples (P1 (338 cm<sup>-1</sup>), P2 (288 cm<sup>-1</sup>), P3 (366 cm<sup>-1</sup>), P4 (376 cm<sup>-1</sup>), P5 (264 cm<sup>-1</sup>), P6 (305–315 cm<sup>-1</sup>) and P7 (412 cm<sup>-1</sup>)).

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