



ELSEVIER

Contents lists available at ScienceDirect

Planetary and Space Science

journal homepage: www.elsevier.com/locate/pss

The science case for an orbital mission to Uranus: Exploring the origins and evolution of ice giant planets



C.S. Arridge^{a,b,*}, N. Achilleos^{c,b}, J. Agarwal^d, C.B. Agnor^e, R. Ambrosi^f, N. André^g, S.V. Badman^{h,i}, K. Baines^{j,k}, D. Banfield^l, M. Barthélémy^{m,br}, M.M. Bisiⁿ, J. Blum^o, T. Bocanegra-Bahamon^p, B. Bonfond^q, C. Bracken^r, P. Brandt^s, C. Briand^t, C. Briosis^u, S. Brooks^j, J. Castillo-Rogez^j, T. Cavalié^v, B. Christophe^w, A.J. Coates^{a,b}, G. Collinson^x, J.F. Cooper^x, M. Costa-Sitja^y, R. Courtin^t, I.A. Dalgis^z, I. de Pater^{aa}, M. Desai^{ab}, D. Dirkx^p, M.K. Dougherty^{ac}, R.W. Ebert^{ab}, G. Filacchione^{ad}, L.N. Fletcher^{ae}, J. Fortney^{af}, I. Gerth^p, D. Grassi^{ad}, D. Grodent^q, E. Grün^{ag,ah}, J. Gustin^q, M. Hedman^{ai}, R. Helled^{aj}, P. Henri^u, S. Hess^{ak}, J.K. Hillier^{al}, M.H. Hofstadter^j, R. Holme^{am}, M. Horanyi^{ah}, G. Hospodarsky^{an}, S. Hsu^{ah}, P. Irwin^{ae}, C.M. Jackman^{ao}, O. Karatekin^{ap}, S. Kempf^{ah}, E. Khalisi^{aq}, K. Konstantinidis^{ar}, H. Krüger^v, W.S. Kurth^{an}, C. Labrianidis^{as}, V. Lainey^{at}, L.L. Lamy^t, M. Laneuville^{au}, D. Lucchesi^{ad}, A. Luntzer^{av}, J. MacArthur^b, A. Maier^{aw}, A. Masters^{h,ac}, S. McKenna-Lawlor^{ax}, H. Melin^f, A. Milillo^{ad}, G. Moragas-Klostermeyer^{aq}, A. Morschhauser^{ay}, J.I. Moses^{az}, O. Mousis^{ba}, N. Nettelmann^{af}, F.M. Neubauer^{bb}, T. Nordheim^{a,b}, B. Noyelles^{bc}, G.S. Orton^j, M. Owens^{bd}, R. Peron^{ad}, C. Plainaki^{ad}, F. Postberg^{aq}, N. Rambaux^{be,at}, K. Retherford^{ab}, S. Reynaud^{bf}, E. Roussos^v, C.T. Russell^{bg}, A.M. Rymer^s, R. Sallantin^g, A. Sánchez-Lavega^{bh}, O. Santolík^{bi}, J. Saur^{bb}, K.M. Sayanagi^{bj}, P. Schenk^{bk}, J. Schubert^{bl}, N. Sergis^{bm}, E.C. Sittler^x, A. Smith^a, F. Spahn^{bn}, R. Srama^{aq}, T. Stallard^{bo}, V. Sterken^{ag,bq}, Z. Sternovsky^{ah}, M. Tiscareno^l, G. Tobie^{bp}, F. Tosi^{ad}, M. Trieloff^{al}, D. Turrini^{ad}, E.P. Turtle^s, S. Vinatier^t, R. Wilson^{ah}, P. Zarka^t

^a Mullard Space Science Laboratory, University College London, UK

^b The Centre for Planetary Science at UCL/Birkbeck, London, UK

^c Department of Physics and Astronomy, University College London, UK

^d ESTEC, European Space Agency, The Netherlands

^e Queen Mary University of London, UK

^f Space Research Centre, University of Leicester, UK

^g IRAP, Toulouse, France

^h ISAS, JAXA, Japan

ⁱ Department of Physics, Lancaster University, UK

^j NASA Jet Propulsion Laboratory, USA

^k University of Wisconsin-Madison, USA

^l Cornell, USA

^m Univ. Grenoble Alpes, IPAG, F-38000 Grenoble, France

ⁿ Rutherford Appleton Laboratory, STFC, UK

^o Technical University, Braunschweig, Germany

^p Delft University of Technology, The Netherlands

^q Université de Liège, Belgium

^r National University of Ireland, Maynooth, Ireland

^s Johns Hopkins University Applied Physics Laboratory, USA

^t LESIA, L'Observatoire de Paris, France

^u LPC2E, CNRS, Université d'Orléans, Orléans, France

^v Max Planck Institute for Solar System Research, Göttingen, Germany

^w ONERA, France

^x NASA Goddard Space Flight Centre, USA

^y ESAC, European Space Agency, The Netherlands

^z Department of Physics, University of Athens, Greece

^{aa} University of California, Berkeley, USA

^{ab} Southwest Research Institute, San Antonio, TX, USA

^{ac} Department of Physics, Imperial College London, UK

- ^{ad} INAF-IAPS Istituto di Astrofisica e Planetologia Spaziali, Rome, Italy
^{ae} Department of Physics, University of Oxford, UK
^{af} University of California Santa Cruz, USA
^{ag} Max Planck Institute for Nuclear Physics, Germany
^{ah} LASP, University of Colorado, USA
^{ai} University of Idaho, Moscow, ID, USA
^{aj} Tel Aviv University, Tel Aviv, Israel
^{ak} LATMOS, France
^{al} Heidelberg University, Germany
^{am} University of Liverpool, UK
^{an} University of Iowa, USA
^{ao} Department of Physics and Astronomy, University of Southampton, UK
^{ap} Royal Observatory of Belgium, Belgium
^{aq} University of Stuttgart, Germany
^{ar} Universität der Bundeswehr München, Germany
^{as} Utesat-Spacecom GmbH, Germany
^{at} IMCCE-Observatoire de Paris, UMR 8028 du CNRS, UPMC, Université Lille 1, 77 Av. Denfert-Rochereau, 75014 Paris, France
^{au} Institut de Physique du Globe de Paris, France
^{av} University of Vienna, Austria
^{aw} Space Research Institute, Austrian Academy of Science, Austria
^{ax} Space Technology Ireland, National University of Ireland, Ireland
^{ay} DLR, Germany
^{az} Space Science Institute, USA
^{ba} Observatoire de Besançon, France
^{bb} University of Cologne, Germany
^{bc} University of Namur, Belgium
^{bd} University of Reading, UK
^{be} Université Pierre et Marie Curie, UPMC - Paris 06, France
^{bf} Laboratoire Kastler Brossel, CNRS, UPMC, France
^{bg} Institute of Geophysics and Planetary Physics, University of California, Los Angeles, USA
^{bh} University of the Basque Country, Spain
^{bi} Institute of Atmospheric Physics, Prague, Czech Republic
^{bj} Department of Atmospheric and Planetary Physics, Hampton University, Virginia, USA
^{bk} Lunar and Planetary Institute, University of Arizona, USA
^{bl} Department of Earth Sciences, University of California, Los Angeles, USA
^{bm} Office for Space Research and Technology, Academy of Athens, Greece
^{bn} University of Potsdam, Germany
^{bo} Department of Physics and Astronomy, University of Leicester, UK
^{bp} LPG, CNRS – Université de Nantes, France
^{bq} International Space Science Institute, Bern, Switzerland
^{br} CNRS, IPAG, F-38000 Grenoble, France

ARTICLE INFO

Article history:

Received 17 December 2013

Received in revised form

23 July 2014

Accepted 7 August 2014

Available online 22 August 2014

Keywords:

Uranus

Magnetosphere

Atmosphere

Natural satellites

Rings

Planetary interior

ABSTRACT

Giant planets helped to shape the conditions we see in the Solar System today and they account for more than 99% of the mass of the Sun's planetary system. They can be subdivided into the Ice Giants (Uranus and Neptune) and the Gas Giants (Jupiter and Saturn), which differ from each other in a number of fundamental ways. Uranus, in particular is the most challenging to our understanding of planetary formation and evolution, with its large obliquity, low self-luminosity, highly asymmetrical internal field, and puzzling internal structure. Uranus also has a rich planetary system consisting of a system of inner natural satellites and complex ring system, five major natural icy satellites, a system of irregular moons with varied dynamical histories, and a highly asymmetrical magnetosphere. Voyager 2 is the only spacecraft to have explored Uranus, with a flyby in 1986, and no mission is currently planned to this enigmatic system. However, a mission to the uranian system would open a new window on the origin and evolution of the Solar System and would provide crucial information on a wide variety of physicochemical processes in our Solar System. These have clear implications for understanding exoplanetary systems. In this paper we describe the science case for an orbital mission to Uranus with an atmospheric entry probe to sample the composition and atmospheric physics in Uranus' atmosphere. The characteristics of such an orbiter and a strawman scientific payload are described and we discuss the technical challenges for such a mission. This paper is based on a white paper submitted to the European Space Agency's call for science themes for its large-class mission programme in 2013.

© 2014 Published by Elsevier Ltd.

1. Introduction

Giant planets account for more than 99% of the mass of the Sun's planetary system, and helped to shape the conditions we see in the Solar System today. The Ice Giants (Uranus and Neptune) are fundamentally different from the Gas Giants (Jupiter and Saturn) in a number of ways and Uranus in particular is the most challenging to our understanding of planetary formation and

* Corresponding author at: Mullard Space Science Laboratory, University College London, UK. Tel.: +44 1483 204 150.
 E-mail address: c.arridge@ucl.ac.uk (C.S. Arridge).

Download English Version:

<https://daneshyari.com/en/article/1781054>

Download Persian Version:

<https://daneshyari.com/article/1781054>

[Daneshyari.com](https://daneshyari.com)