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Measurement of the half-life for two ⁵⁷Fe excited states by a single-crystal scintillation time spectrometer

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Abstract

Half-lives measurements for two excited states of ⁵⁷Fe located at 14.4 and 136.5 keV, in the ⁵⁷Fe nucleus have been recently proposed using the Autocorrelation single-crystal scintillation time spectrometer (ASCSTS) .We have proved that ASCSTS has a reasonably high detection efficiency for $\gamma - \gamma$ coincidences in the energy range around 5 keV, preserving its main distinctive feature: a large detection solid angle (close to 2π) in a wide time interval (from ~4 ns to 100 µs).

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In a recent work [1] some of us introduced the autocorrelation single-crystal scintillation time spectrometer (ASCSTS) with plastic scintillators, as a new instrument for studying nuclear isomeric states in the nano- and microsecond range. The new experimental method ensures detection with high efficiency of the delayed $\gamma - \gamma$ coincidences, in a geometry closed to 4π . It allows the measurement of the half-lives for excited and ground states, in range of 4 ns to 100 µs. The use of a few millimeters thick plastic scintillators with short decay times, provides almost 100% detection efficiency of the charged particles (conversion electrons, β^{-} and α particles) emitted in radioactive decays. Generally, isomeric states are populated or depopulated by low-energy γ transitions, whose internal conversion coefficients are rather large. The low-energy limit for detecting radiation is defined only by the electronic noise of the (PMT). In our experiments we investigated very weak intensity sources (the activity of the studied sources was comparable with the background level

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in the laboratory) with complex isotopic composition resulting from long radioactive chains beginning with ²²⁵Ac (8 radioactive decays) and ²³²Th (11 radioactive decays) by measuring β^- -e⁻, α -e⁻ or β^- - α delayed coincidences [1,2].

In the present paper, we attempt to test our ASCSTS instrument, working with optimum performance in the regime of γ -e⁻ and/or γ - γ delayed coincidences, using it to study nuclear isomeric states in ⁵⁷Fe. ⁵⁷Co decays to ⁵⁷Fe which has two short-lived isomeric states, at 14.4 and 136.5 keV. The ⁵⁷Co radioactive source was obtained by irradiating a natural iron foil (4 µm thickness) with a proton beam at the cyclotron accelerator at NIPNE-HH. The partial scheme of the ⁵⁷Co decay is shown in Fig. 1, [3].

The block diagram of the single-crystal time spectrometer is presented in Fig. 2. A XP2020 PMT, with low noise level ($\leq 5 \text{ keV}$) and a plastic scintillator type NE104 (40 mm diameter and 20 mm length) and decay time of $\tau = 1.9$ ns have been used. Start-and stop-time signals are taken from the constant fraction discriminator (D, ORTEC 473). By its design principle, ASCSTS uses only one-time discriminator with a threshold set to a value

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Fig. 1. Fragment of the ⁵⁷Co decay scheme [3].



Fig. 2. Block diagram of the autocorrelation single-crystal time spectrometer. PMT—photomultiplier, FA—fast amplifier, DL—delay line, LIN—linear integration nanosecond pulse circuit, D—discriminator, TR—trigger, CC—coincidence circuit, TAC—time-to-amplitude converter.

slightly exceeding the level of PMT noise. This leads to: (i) the most highly achievable efficiency of recording populating and delayed radiation and (ii) a simple spectrometer with increased reliability. The ultimate stop signals are generated by a gated coincidence circuit (CC). The trigger (TR), produces a pulse with a width equal to the duration of the studied time range. A delay line (DL) eliminates selfcoincidences. The signal taken from the PMT dynode, transferred through a fast amplifier (FA, ORTEC 579) to the linear nanosecond pulse integration circuit (DL), allows two-dimensional energy-time coincidence spectra. Decay scheme of ⁵⁷Co is very well known (the decay process populates only the two interesting levels: 14.4 and 136.5 keV), we did not use in the analysis the energy information and we have sorted the experimental data using only the time distribution spectra. The scintillation detector cannot ensure a satisfactory energy analysis due to poor energy resolution of the plastic scintillator. The precise determination of the energy for the isomeric state can be done by adding to the experimental setup a highpurity germanium detector [2].

In the present experiment the energy range was selected above $E \sim 5 \text{ keV}$. The measurements were performed in a geometry closed to 2π . The ⁵⁷Co radioactive source was placed at 1 mm from the face of the plastic scintillator. The conversion electrons from the 14.4 keV energy transition were completely absorbed by a polyethylene filter placed between the source and the plastic detector [4]. In this way, we reproduced the experimental conditions for the ASCSTS, similar to an on-line experiment with a thick target.

The results of the measurements are shown in Fig. 3. Curves 1 and 2 correspond to the time distribution of ⁵⁷Co decays and curve 3 to ¹³⁷Cs case. The main difference between the curves 1 and 2 is caused by the use of a thick polyethylene foil (1 mm thick) filter. Although a strong absorption of γ -rays having 14.4 keV energy takes place (the absorption coefficient equal to three), both curves have the same shape, but in case 2 the counting rate was lower. Curve 3 represents the time spectrum recorded with a radioactive ¹³⁷Cs source which gives an estimate of the time distribution for the accidental coincidences. All these spectra were recorded under the same working conditions.

In order to extract the half-life values from the experimental data, time scale calibration for the time spectrometer and an appropriate implementation of unfolding procedure for delayed coincidence spectra are required. A time calibration for the time-to-amplitude converter (TAC) has been obtained by introducing



Fig. 3. Time spectra at decay of 57 Co (curves 1 and 2) and 137 Cs (curve 3). For the clarity of the figure, the amount of data in each channel was multiplied by 0.5 for the curve number 2 and by 0.1 for curve number 3.

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