

Fermionic extension of the scalar Born–Infeld equation and its relation to the supersymmetric Chaplygin gas

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Abstract

In this Letter, we present a fermionic extension of the scalar Born–Infeld equation, which has been derived from the Nambu–Goto superstring action in $(2 + 1)$ dimensions through the Cartesian parameterization. It is demonstrated that in the relativistic limit where $c \rightarrow \infty$, the fermionic Born–Infeld model reduces to the supersymmetric Chaplygin gas model in one spatial dimension. However, the supersymmetry itself is not preserved.

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1. Introduction

The theory of nonlinear electrodynamics of Born and Infeld had its origins in the work of Mie [1], who attempted to formulate a unitarian theory of electromagnetics through a nonlinear generalization of Maxwell's equations. However, Mie's theory possessed the disadvantage that the electromagnetic potentials acquired a direct physical significance, leading to the loss of gauge invariance. Born and Infeld postulated a nonlinear, gauge-invariant, relativistic theory [2,3] described by the Lagrangian

$$L_{\text{BI}} = b^2 \left(1 - \sqrt{1 - (\mathbf{E}^2 - \mathbf{B}^2)/b^2 - (\mathbf{E} \cdot \mathbf{B})^2/b^4} \right), \quad (1)$$

where \mathbf{E} and \mathbf{B} represent the electric and magnetic fields respectively, and b is a parameter having the dimension of the electromagnetic field [3,4]. The laws of propagation of photons and charged particles were studied by Boillat [5], and it was shown that the Born–Infeld theory leads to propagation with-

out birefringence or shock waves. More generally, equations derived from an action principle involving a square root similar to that of (1) are referred to as Born–Infeld type equations in the literature on the subject. Such generalized Born–Infeld Lagrangians have appeared, for example, in string theories involving scalar (dilaton) fields and gravity [6,7]. More recently, a generalization of the Born–Infeld Lagrangian for non-Abelian gauge theory was proposed [8] and adapted to the noncommutative geometry of matrix valued functions on a manifold [9].

A few years ago, a series of lectures was given by Professor R. Jackiw in which the subject of fluid mechanics was examined from an entirely new perspective [10]. A number of topics was covered, including the equations of motion, conservation laws, and a general description of the properties of classical fluids. Of particular interest were the Galileo-invariant Chaplygin gas model, which describes an isentropic, non-dissipative fluid with polytropic pressure [11,12], as well as the Poincaré-invariant Born–Infeld scalar model used to describe the interaction of two plane waves [13,14]. In particular, it was shown that both models devolve from the parameterization-invariant Nambu–Goto action for a d -brane evolving in $(d + 1, 1)$ -dimensional space–

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time

$$I_{\text{NG}} = \int \sqrt{(-1)^d \det\left(\frac{\partial X^\mu}{\partial \phi^i} \frac{\partial X_\mu}{\partial \phi^j}\right)} d\phi^0 d\phi^1 \dots d\phi^d. \quad (2)$$

Here, $\phi^0, \phi^1, \dots, \phi^d$ are the world-volume variables which parameterize the extended object, while the variables $X^0, X^1, \dots, X^d, X^{d+1}$ describe the target space–time. Two distinct choices of parameterization of the action (2) lead respectively to the Chaplygin and Born–Infeld models. In both cases, (X^1, \dots, X^d) is chosen to coincide with (ϕ^1, \dots, ϕ^d) , and they are called \mathbf{r} (a d -dimensional position vector). For the light-cone parameterization, the quantity $(X^0 + X^{d+1})/\sqrt{2}$ is renamed t and identified with $\sqrt{2\lambda}\phi^0$, while $(X^0 - X^{d+1})/\sqrt{2}$ is identified with θ , a function of t and \mathbf{r} . The Nambu–Goto action (2) reduces to the Chaplygin gas action

$$I_\lambda = -2\sqrt{\lambda} \int \sqrt{\partial_t \theta + \frac{1}{2}(\nabla \theta)^2} dt d\mathbf{r}. \quad (3)$$

For the Cartesian parameterization, X^0 is renamed ct and identified with ϕ^0 , while X^{d+1} is renamed θ/c , a function of t and \mathbf{r} . The Nambu–Goto action (2) then coincides with the scalar Born–Infeld action

$$I_a = -a \int \sqrt{c^2 - (\partial_\mu \theta)^2} dt d\mathbf{r}. \quad (4)$$

It should also be noted that at the nonrelativistic limit where $c \rightarrow \infty$, the Born–Infeld action (4) reduces to the Chaplygin action (3) provided that the nonrelativistic function θ_{NR} in (3) is extracted from the relativistic function θ_{R} in (4) by the relation

$$\theta_{\text{R}} = -c^2 t + \theta_{\text{NR}}. \quad (5)$$

Moreover, it was shown that fluid mechanics can be enhanced through the addition of fermionic (anticommuting Grassmannian) degrees of freedom. In particular, supersymmetric generalizations were formulated for the Chaplygin gas in one and two spatial dimensions [15,16]. In the case of one spatial dimension, the resulting Lagrangian density reads

$$\mathcal{L} = -\sqrt{2\partial_t \theta - \psi \partial_t \psi + \left(\partial_x \theta - \frac{1}{2}\psi \partial_x \psi\right)^2} + \frac{1}{2}\psi \partial_x \psi, \quad (6)$$

where θ is the standard bosonic scalar field of the Chaplygin gas equation and ψ is a real fermionic Grassmann variable.

It was also demonstrated how the equations for the supersymmetric Chaplygin fluid devolve from a supermembrane Lagrangian, through the light-cone parameterization mentioned above. The question arises as to whether an analogous supersymmetric extension can be formulated for the scalar Born–Infeld model by applying the Cartesian parameterization to the supermembrane Lagrangian. In this Letter, we attempt to answer this question for the case of one spatial dimension.

This Letter is organized as follows. In Section 2, we parameterize the Nambu–Goto action for a superstring using the Cartesian parameterization and evaluate the resulting Lagrangian. It is demonstrated that this Lagrangian is indeed a Lorentz-invariant fermionic extension of the one for the scalar Born–Infeld equation. Furthermore, in the nonrelativistic limit where

$c \rightarrow \infty$, the theory reduces to that of the supersymmetric Chaplygin Lagrangian in one spatial dimension (6). In Section 3, we discuss the Hamiltonian and the canonical form of the Lagrangian and derive the equations of motion. Finally in Section 4, we discuss a supersymmetry of the Nambu–Goto superstring theory and the fact that this supersymmetry is not carried over when we go to the Cartesian parameterization.

2. Fermionic extension of the scalar Born–Infeld model

We begin with the Nambu–Goto action for a superstring in a $(2+1)$ -dimensional target space–time

$$I = - \int (\sqrt{g} - i\epsilon^{ij} \partial_i X^\mu \bar{\psi} \gamma_\mu \partial_j \psi) d\phi^0 d\phi^1, \quad (7)$$

where

$$g = -\det[(\partial_i X^\mu - i\bar{\psi} \gamma^\mu \partial_i \psi)(\partial_j X^\nu - i\bar{\psi} \gamma^\nu \partial_j \psi) \eta_{\mu\nu}]. \quad (8)$$

Henceforth, μ and ν are indices running over 0, 1, 2, which represent the target space–time, and i, j are the worldsheet indices 0, 1. We use the symbol ∂_i to denote derivation by ϕ^i and the γ^μ matrices are defined as

$$\gamma^0 = \sigma^1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \gamma^1 = i\sigma^2 = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix},$$

$$\gamma^2 = i\sigma^3 = \begin{pmatrix} i & 0 \\ 0 & -i \end{pmatrix}, \quad (9)$$

$$\gamma^5 = \gamma^0 \gamma^1 = \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix} = i\gamma^2. \quad (10)$$

Here, ψ is a real two-component spinor $\psi = (u, v)$ whose components are odd (fermionic) Grassmann variables. For the purpose of deriving the generalized fermionic Born–Infeld action, we chose the following Cartesian parameterization

$$\begin{aligned} X^0 &= ct, & X^1 &= x, & X^2 &= \frac{1}{c}\theta(t, x), \\ \phi^0 &= ct, & \phi^1 &= x. \end{aligned} \quad (11)$$

The components of the matrix g_{ij} in Eq. (8) are then given by

$$\begin{aligned} g_{ij} &= \eta_{ij} - \frac{1}{c^2}(\partial_i \theta)(\partial_j \theta) - i\bar{\psi} \gamma_i \partial_j \psi - i\bar{\psi} \gamma_j \partial_i \psi \\ &\quad + \frac{1}{c}i(\partial_i \theta)\bar{\psi} \gamma^2 \partial_j \psi + \frac{1}{c}i(\partial_j \theta)\bar{\psi} \gamma^2 \partial_i \psi + 3\bar{\psi} \partial_i \psi \bar{\psi} \partial_j \psi, \end{aligned} \quad (12)$$

where we have used the identity

$$(\sigma^\mu)_{ij}(\sigma^\mu)_{kl} = 2\delta_{il}\delta_{jk} - \delta_{ij}\delta_{kl} \quad (\mu = 0, 1, 2). \quad (13)$$

The supplementary term in the action (7) is

$$-i\epsilon^{ij} \bar{\psi} \gamma_i \partial_j \psi + \frac{1}{c}i\epsilon^{ij}(\partial_i \theta)\bar{\psi} \gamma^2 \partial_j \psi. \quad (14)$$

The fermionic gauge choice

$$(1 + \gamma^5)\psi = 0, \quad (15)$$

reduces the spinor ψ to a one-component odd Grassmann field, which we renormalize to $\frac{1}{\sqrt{2ic}}u$, with real u . The components

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