

# Aspects of light propagation in anisotropic dielectric media

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## Abstract

Some aspects of light propagation in local anisotropic nonlinear dielectric media at rest in the limit of geometrical optics are investigated. Natural and artificially induced anisotropies in dielectric materials are discussed. Analogies are proposed in such way that, as far as light is considered, kinematic aspects of some cosmological models are recovered. Particularly, analogue models for isotropic and anisotropic cosmologies are presented.

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## 1. Introduction

Inside material media electrodynamics becomes nonlinear. In such situations the Maxwell equations must be supplemented with constitutive relations which, in general, are nonlinear and depend on the physical properties of the medium under the action of external fields. As a consequence, several effects (unusual in the context of linear Maxwell theory) are predicted. Of actual interest is the phenomenon of artificial birefringence: when an external field is applied in a medium with nonlinear dielectric properties, an artificial optical axis may appear [1–5].

The development of analogies in order to test kinematic aspects of general relativity in laboratory have been performed in several branches of physics [2,3,5–12]. Particularly, nonlinear electrodynamics has been considered as a possible scenario to construct analogue models for general relativity, either in the context of nonlinear Lagrangian or nonlinear material media. This is based in the fact that the trajectory of photons can be described by a null geodesic in an effective metric  $g_{\mu\nu}$ . In this work, homogeneous dielectric media at rest with the dielectric

coefficients  $\varepsilon^\mu_\nu(\vec{E})$  and constant  $\mu$ , in the limit of geometrical optics, are used to construct analogue models for cosmology. The analysis is restricted to local electrodynamics. In order to avoid ambiguities with the wave velocity, dispersive effects were neglected by considering only monochromatic waves. It is shown that naturally uniaxial media presenting nonlinear dielectric properties can be operated by external fields in such way to induce anisotropy in the optical metric.

A covariant formalism is used throughout this work. Space-time is assumed to be Minkowskian, and a Cartesian coordinate system is used, such that the metric is  $\eta_{\mu\nu} = \text{diag}(+1, -1, -1, -1)$ . Units are chosen such that  $c = 1$ . A geodetic observer  $V^\mu = \delta^\mu_0$  is supposed to describe all quantities. Particularly the electric field is represented by  $E^\mu = -F^{\mu\nu}V_\nu = (0, \vec{E})$  whose modulus is  $E = (-E^\alpha E_\alpha)^{1/2}$ .

## 2. The dispersion relation

The properties of light propagation in material media are determined by the so-called dispersion relations, which can be derived, in the context of the eikonal approximation of electrodynamics, making use of the method of field discontinuities [13]. Define a surface of discontinuity  $\Sigma$  by  $z(x^\alpha, \vec{x}) = 0$ . Whenever  $\Sigma$  is an inextendible surface, it divides the spacetime in two disjoint regions  $U^-$  for  $z(x^\alpha, \vec{x}) < 0$ , and  $U^+$  for  $z(x^\alpha, \vec{x}) > 0$ .

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The discontinuity of an arbitrary function  $f(x^0, \vec{x})$  on  $\Sigma$  is given by

$$[f(x^0, \vec{x})]_{\Sigma} \doteq \lim_{\{P^{\pm}\} \rightarrow P} [f(P^+) - f(P^-)] \quad (1)$$

with  $P^+$ ,  $P^-$  and  $P$  belonging to  $U^+$ ,  $U^-$  and  $\Sigma$ , respectively. The electric and magnetic fields are continuous when crossing the surface  $\Sigma$ . However, their derivatives behave as

$$[E^{\mu}, v]_{\Sigma} = e^{\mu} K_{\nu}; \quad [B^{\mu}, v]_{\Sigma} = b^{\mu} K_{\nu}, \quad (2)$$

where  $e^{\mu}$  and  $b^{\mu}$  represent the discontinuities of the fields on the surface  $\Sigma$  and

$$K_{\lambda} = \frac{\partial \Sigma}{\partial x^{\lambda}} \quad (3)$$

is the wave vector.

When these conditions are applied to the electrodynamic field equations in local anisotropic dielectric media, the following dispersion relation for light rays [5] are obtained:

$$g_{\pm}^{\lambda\tau} K_{\lambda} K_{\tau} = \left\{ \mu\alpha V^{\lambda} V^{\tau} + \frac{1}{2} \left[ C^{\nu}_{\nu} - \frac{1}{\mu(v_{\varphi}^{\pm})^2} \right] C^{(\lambda\tau)} - \frac{1}{2} C^{(\lambda}_{\nu} C^{\nu\tau)} \right\} K_{\lambda} K_{\tau} = 0, \quad (4)$$

where

$$C^{\alpha}_{\tau} \doteq \varepsilon^{\alpha}_{\tau} + \frac{\partial \varepsilon^{\alpha}_{\beta}}{\partial E^{\tau}} E^{\beta} + \frac{1}{\omega} \frac{\partial \varepsilon^{\alpha}_{\beta}}{\partial B^{\rho}} \eta^{\rho\lambda\gamma}_{\tau} E^{\beta} K_{\lambda} V_{\gamma} \quad (5)$$

and the phase velocities  $v_{\varphi}^{\pm}$  are

$$v_{\varphi}^{\pm} = \sqrt{\frac{\beta}{2\alpha} \left( 1 \pm \sqrt{1 + \frac{4\alpha\gamma}{\beta^2}} \right)}, \quad (6)$$

with  $\omega \doteq K^{\alpha} V_{\alpha}$  the frequency of the electromagnetic wave and the coefficients  $\alpha$ ,  $\beta$  and  $\gamma$  given by

$$\alpha \doteq \frac{1}{6} [(C^{\mu}_{\mu})^3 - 3C^{\mu}_{\mu} C^{\alpha}_{\beta} C^{\beta}_{\alpha} + 2C^{\alpha}_{\beta} C^{\beta}_{\gamma} C^{\gamma}_{\alpha}], \quad (7)$$

$$\beta \doteq \mu^{-1} (C^{\lambda}_{\alpha} C^{\alpha\nu} - C^{\alpha}_{\alpha} C^{\lambda\nu}) \hat{q}_{\lambda} \hat{q}_{\nu}, \quad (8)$$

$$\gamma \doteq \mu^{-2} C^{\lambda\nu} \hat{q}_{\lambda} \hat{q}_{\nu}. \quad (9)$$

In the last two Eqs. (8)–(9) we introduced the 3-dimensional projection of the wave vector  $K^{\alpha}$  as  $q^{\alpha} = h^{\alpha}_{\mu} K^{\mu} = K^{\alpha} - \omega V^{\alpha}$ , and  $\hat{q}^{\mu} = q^{\mu}/q$ . We defined the projector  $h^{\alpha}_{\mu} = \delta^{\alpha}_{\mu} - V^{\alpha} V_{\mu}$ .

The symmetric tensors  $g_{\pm}^{\mu\nu}$  are the optical coefficients<sup>1</sup> associated with the wave propagation, and the symbol  $\pm$  indicates the possibility of two distinct coefficients, one for each polarization mode—birefringence phenomena. Correspondingly to it Eq. (6) expresses the fact that, in general, the phase velocity of the electromagnetic waves inside a material medium may get two possible values ( $v_{\varphi}^{+}$ ,  $v_{\varphi}^{-}$ ) which are associated with the two possible polarization modes. For the particular case of Maxwell

linear theory in vacuum, both  $g_{+}^{\mu\nu}$  and  $g_{-}^{\mu\nu}$  reduce to the diagonal matrix  $(+1, -1, -1, -1)$ , which is identified with the Minkowski metric  $\eta^{\mu\nu}$ , as expected. Other particular cases are obtained by considering isotropic media [2,3,7] and some applications was recently proposed in the context of dielectric analogues of black hole spacetime [9].

### 3. Naturally anisotropic uniaxial media

Now, let us consider naturally anisotropic uniaxial media reacting nonlinearly when subjected to an external electric field as  $\varepsilon^{\alpha}_{\beta} = \text{diag}[0, \varepsilon_{\parallel}(E), \varepsilon_{\perp}(E), \varepsilon_{\perp}(E)]$ . In this case  $\varepsilon^{\alpha}_{\beta} = \varepsilon^{\alpha}_{\beta}(E)$  and by setting  $\vec{E}$  in the  $x$ -direction (optical axis) we obtain  $C^{\alpha}_{\beta} = \text{diag}(0, \varepsilon_{\parallel} + E\varepsilon'_{\parallel}, \varepsilon_{\perp}, \varepsilon_{\perp})$ , where  $\varepsilon'_{\parallel} = d\varepsilon_{\parallel}/dE$ . For this particular case  $C^{\alpha\beta}$  is a symmetric tensor. The phase velocities reduce to

$$(v_{\varphi}^{+})^2 = \frac{1}{\mu\varepsilon_{\perp}}, \quad (10)$$

$$(v_{\varphi}^{-})^2 = \frac{1}{\mu\varepsilon_{\perp} C^1_1} [\varepsilon_{\perp} (1 - \hat{q}_1^2) + C^1_1 \hat{q}_1^2]. \quad (11)$$

Note that  $v_{\varphi}^{-}$  depends on the direction of propagation, as it should be expected for the extraordinary ray. The two velocities coincide when either the propagation occurs along the direction of the electric field ( $\hat{q}_1^2 = 1$ ), or when the no-birefringence condition  $\varepsilon_{\parallel} + E\varepsilon'_{\parallel} = 0$  holds [5].

Let us also particularize to the model where

$$\varepsilon_{\perp} = \varepsilon_{\perp} - 3pE^2, \quad (12)$$

$$\varepsilon_{\parallel} = \varepsilon_{\parallel} - sE^2. \quad (13)$$

Thus,  $C^{\alpha}_{\beta} = \text{diag}(0, \varepsilon_{\parallel} - 3sE^2, \varepsilon_{\perp} - 3pE^2, \varepsilon_{\perp} - 3pE^2)$ .

For the ordinary ray the optical coefficients are

$$g_{+}^{00} = \mu\alpha, \quad (14)$$

$$g_{+}^{ii} = -\varepsilon_{\parallel}\varepsilon_{\perp} + 3(s\varepsilon_{\perp} + p\varepsilon_{\parallel})E^2 - 9spE^4, \quad (15)$$

where

$$\alpha = -27sp^2E^6 + 9p(p\varepsilon_{\parallel} + 2s\varepsilon_{\perp})E^4 - 3\varepsilon_{\perp}(s\varepsilon_{\perp} + 2p\varepsilon_{\parallel})E^2 + \varepsilon_{\parallel}\varepsilon_{\perp}^2. \quad (16)$$

Eqs. (14)–(15) show that for the ordinary ray there will be no anisotropy in the space section.

For the extraordinary ray the optical coefficients are

$$g_{-}^{00} = \mu\alpha, \quad (17)$$

$$g_{-}^{11} = -(\chi - \varepsilon_{\parallel} + 3sE^2)(\varepsilon_{\parallel} - 3sE^2), \quad (18)$$

$$g_{-}^{22} = g_{-}^{33} = -(\chi - \varepsilon_{\perp} + 3pE^2)(\varepsilon_{\perp} - 3pE^2), \quad (19)$$

where  $\chi$  depends on the direction on wave propagation as

$$\chi = -\frac{(\varepsilon_{\parallel} - 3sE^2)(\varepsilon_{\perp} - 3pE^2)}{(\varepsilon_{\perp} - 3pE^2)(1 - \hat{q}_1^2) + (\varepsilon_{\parallel} - 3sE^2)\hat{q}_1^2} + (\varepsilon_{\parallel} - 3sE^2) + 2(\varepsilon_{\perp} - 3pE^2). \quad (20)$$

Eqs. (17)–(19) show that for the extraordinary ray there will be anisotropy in the section ( $g_{-}^{11} \neq g_{-}^{22} = g_{-}^{33}$ ).

<sup>1</sup> Such coefficients are called in the literature as the components of the optical metric, meanly in the case where they does not depend on the direction of light propagation.

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