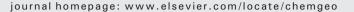
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# Chemical Geology



# Trace element partitioning between perovskite and kimberlite to carbonatite melt: New experimental constraints

# Christopher Beyer<sup>1</sup>, Jasper Berndt, Sebastian Tappe, Stephan Klemme<sup>\*</sup>

Institut für Mineralogie, Westfälische-Wilhelms-Universität Münster, Corrensstrasse 24, 48149 Münster, Germany

## ARTICLE INFO

## ABSTRACT

*Article history:* Accepted 27 March 2012 Available online 4 April 2012

Keywords: Perovskite Partition coefficients Kimberlite Carbonatite Alkaline magma petrogenesis LA-ICP-MS We present experimentally determined trace element partition coefficients (D) between perovskite (CaTiO<sub>3</sub>) and several different silicate and carbonate melt compositions. Our data indicate that perovskite is able to strongly fractionate trace elements during the evolution of SiO<sub>2</sub>-undersaturated magmas. In silicate magmas, perovskite efficiently fractionates Zr and Hf from Nb and Ta, with D<sub>Zr</sub> and D<sub>Hf</sub> coefficients below and D<sub>Nb</sub> and D<sub>Ta</sub>, significantly above unity, respectively. All trace element partition coefficients between perovskite and the diverse silicate melts, which primarily differ in their CaO and SiO<sub>2</sub> contents, are fairly similar within propagated uncertainties. We find that D perovskite/silicate melt for the high field strength elements is much lower than corresponding perovskite/carbonate melt partition coefficients for the same elements. Another key difference between perovskite/melt trace element partitioning in silicate and carbonate systems is the relative behavior of U and Th. Whereas all silicate system experiments show that D<sub>Th</sub> > D<sub>U</sub>, the relationship is different in the carbonatite experiment. This implies that perovskite may fractionate U and Th in silicate perovskite may fractionate U and Th in silicate and carbonate systems. The rare earth elements (REE) show similar perovskite/melt partitioning in the silicate and carbonate

systems investigated. Our data indicate that the REE are highly compatible in perovskite in both silicate and carbonatic compositions, which explains the high rare earth element concentrations in naturally occurring magmatic perovskites. In particular, our new trace element partition coefficients help to better understand and to quantify the strong control of perovskite on the bulk REE budget of kimberlites and related rocks. © 2012 Elsevier B.V. All rights reserved.

1. Introduction

Understanding trace element behavior in magmatic systems requires a thorough understanding of the partitioning of trace elements between minerals and melts. Previous studies have focused on the trace element partitioning between major rock forming minerals and melts (Green, 1994), and relatively few partition coefficients are available for accessory phases such as rutile, ilmenite, apatite, or perovskite (e.g., Nagasawa et al., 1980; Kennedy et al., 1994; Corgne and Wood, 2002; Klemme and Dalpe, 2003; Klemme and Meyer, 2003; Corgne and Wood, 2005; Klemme et al., 2005; Prowatke and Klemme, 2005; Klemme et al., 2006). The phases mentioned above typically do not play an important role during the fractionation of basaltic magmas, but they certainly exert strong control on the trace element distribution in alkaline, kimberlite, and carbonatite magmatic systems where they occur throughout the crystallization sequence and reach abundance levels commonly exceeding 5 vol.%, i.e., they can be minor rock forming constituents (Mitchell, 1986; Tappe et al., 2004; Tappe et al., 2009; Malarkey et al., 2010; Chen and Simonetti, 2013–this issue).

In this study we focus on CaTiO<sub>3</sub>-perovskite, which is a common phase in silica-undersaturated alkaline (Veksler and Teptelev, 1990; Dawson, 1994; Tappe et al., 2005; Rosenthal et al., 2009), kimberlite (Mitchell, 1986; Chakhmouradian and Mitchell, 2000; Malarkey et al., 2010; Rao et al., 2011; Tappe et al., 2011) and, more rarely, carbonatite magmas (Chakhmouradian and Mitchell, 1997; Chakhmouradian, 2006; Cox and Wilton, 2006). Although perovskites from different alkaline and kimberlite varieties appear to be close to CaTiO<sub>3</sub> end-member composition (Mitchell, 1995; Chakhmouradian and Mitchell, 2000; Tappe et al., 2006), perovskites from carbonatites show relatively large compositional diversity, in particular in terms of their REE and Nb contents (Lupini et al., 1992; Platt, 1994; Viladkar and Subramanian, 1995; Mitchell, 2002; Chakhmouradian, 2006; Reguir et al., 2010). Whether this behavior is due to the typically elevated REE and Nb concentrations of carbonaterich magmas, or simply a function of distinctly different perovskite/melt trace element partitioning between silicate and carbonate magmatic





<sup>\*</sup> Corresponding author. Tel.: +49 251 8333047; fax: +49 251 8338397.

E-mail addresses: christopher.beyer@uni-bayreuth.de (C. Beyer),

jberndt@uni-muenster.de (J. Berndt), sebastian.tappe@uni-muenster.de (S. Tappe), stephan.klemme@uni-muenster.de (S. Klemme).

<sup>&</sup>lt;sup>1</sup> Present address: Bayerisches Geoinstitut, University of Bayreuth, Germany.

<sup>0009-2541/\$ –</sup> see front matter 0 2012 Elsevier B.V. All rights reserved. doi:10.1016/j.chemgeo.2012.03.025

systems, will be addressed during the discussion of our new experimental results.

Corgne and Wood (2002, 2005) investigated the partitioning of trace elements at varying P-T conditions in the system SiO<sub>2</sub>-CaO- $TiO_2 \pm Al_2O_3$ . Their data suggest that pressure and temperature have only minor influence on the partitioning of trace elements between CaTiO<sub>3</sub> and silicate melts. However, they observed a change in the partitioning behavior with varying CaO contents of the melt, suggesting that melt composition has a strong effect on trace element partitioning (Kohn and Schofield, 1994; Klemme, 2004; Klemme et al., 2005; Prowatke and Klemme, 2005, 2006a,b; Schmidt et al., 2006; Evans et al., 2008). The objective of our study is to extend the database to more natural bulk compositions, and to evaluate the applicability of simple system partition coefficients to more complex silicate and carbonate melt compositions. Therefore, we initially determined perovskite/melt trace element partition coefficients in a simple system following Corgne and Wood (2005) and compared these data with our results from a more complex H<sub>2</sub>O- and CO<sub>2</sub>bearing silicate system, the average near-primary kimberlite composition (Becker and Le Roex, 2006). In addition, we performed an experiment in a simplified carbonatitic system, which enables us to present the first perovskite/carbonatite melt trace element partitioning data.

#### 2. Starting material and methods

### 2.1. Starting mixtures and experimental approach

To ensure the nucleation and growth of perovskite to a reasonable size and, thus, enabling in situ analysis by laser ablation inductivelycoupled plasma mass spectrometry (LA-ICP-MS), the starting materials were super-saturated in CaO and TiO<sub>2</sub>. In order to avoid possible experimental problems related to iron-loss, the compositions were all iron-free. Mix 1 is a composition in the system SiO<sub>2</sub>-CaO-TiO<sub>2</sub>, whereas Mix 2 is more complex in the system SiO<sub>2</sub>-CaO-TiO<sub>2</sub>-MgO-K<sub>2</sub>O-Al<sub>2</sub>O<sub>3</sub>. All compositions were prepared from analytical grade oxides, hydroxides, and carbonates (Table 1), which were ground in an agate mortar under ethanol. For the high-pressure partitioning experiments in a kimberlitic system, the composition of our starting material (MixBeyK1) was guided by an average near-primary kimberlite composition from southern Africa (Becker and Le Roex, 2006). Furthermore, a model carbonatite composition was prepared in the system CaCO<sub>3</sub>-TiO<sub>2</sub> (SKHDCMB2). The choice of this rather simple composition was made after our reconnaissance experiments in more complex carbonate systems yielded only small perovskite crystals  $(<15 \,\mu\text{m})$ , which were impossible to analyze by our LA-ICP-MS set-up.

Mix 1 and Mix 2 were heated in air from 800 °C to 1000 °C with a ramp of 150 °C/h. They were then held at 1000 °C for 3 h in order to ensure complete decarbonation. Subsequently, the mixtures were heated to temperatures well above the liquidus, quenched, and again

Table 1	
Starting material compositions	(wt.%).

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		Mix 1	Mix 2	MixBeyK1	SKHDCMB2
	SiO <sub>2</sub>	17.5	25	28	
	TiO <sub>2</sub>	54	20	11.2	27.5
	CaO	28.5	15	17	40.6
	$Al_2O_3$		2.5	2.7	
	MgO		20	26	
	NiO			0.1	
	Na <sub>2</sub> O				
	K <sub>2</sub> O		2.0	1.0	
	H <sub>2</sub> O			8	
	CO <sub>2</sub>			6	31.9

ground to fine powders. MixBeyK1 was decarbonated in the same manner, but  $Mg(OH)_2$  and CaCO<sub>3</sub> were added at a later stage as sources of  $H_2O$  and  $CO_2$ , respectively. For the carbonatite starting material SKHDCMB2, CaCO<sub>3</sub> and TiO<sub>2</sub> were mixed in the appropriate ratios. All starting materials were doped with a trace element solution containing the following elements in 2% HNO<sub>3</sub> solution: Li, Sr, Ba, Sc, Zr, Nb, La, Ce, Pr, Sm, Gd, Lu, Y, Hf, Ta, Pb, U, and Th. This solution was prepared from 1000  $\mu g/g$  ICP-AES standard solutions. The doped starting materials were dried and denitrified in a laboratory-type drying cabinet at 110 °C overnight.

The 1-atmosphere experiments with Mix 1 and Mix 2 as starting materials were carried out in two different gas-mixing furnaces under air. The temperature was monitored and controlled with a Type B thermocouple ( $Pt_{70}Rh_{30}$ – $Pt_{94}Rh_6$ ). The runtime varied between 18 and 66 h (see Table 2 for details). After quenching of the experimental charges, the samples were embedded in epoxy resin and polished using a series of diamond pastes.

The high-pressure experiments were performed in an end-loaded piston cylinder apparatus (Boyd and England, 1960) at 1.5 GPa (Table 2). This pressure is representative for the uppermost mantle where the first phenocrysts of alkaline magmas are thought to crystallize (Mitchell, 1986). The samples were enclosed in a 2 mm O.D. platinum capsule. The pressure assemblies consisted of a 1/2 inch talc-Pyrex tube, and a 6 mm diameter graphite heater. Cavities were filled with crushable alumina ceramics. The temperature was monitored by  $W_{97}Re_3-W_{75}Re_{25}$  thermocouples. A friction correction of -13% was applied which was calibrated for this particular assembly on the quartz-coesite transition (Bose and Ganguly, 1995) and the MgCr<sub>2</sub>O<sub>4</sub> + SiO<sub>2</sub> = MgSiO<sub>3</sub> + Cr<sub>2</sub>O<sub>3</sub> reaction (Klemme and O'Neill, 1997). Run durations varied between 5 and 8 h. Based on our pressure calibration, the quoted pressures are estimated to be accurate within 0.1 GPa (Table 2).

### 2.2. Analytical methods

Run products were examined qualitatively using a JEOL 6610LV scanning electron microscope with EDX system. Major element analyses were performed with a 4-spectrometer JEOL JXA 8900 electron microprobe at the University of Münster using an acceleration voltage of 15 kV. Mineral phases were analyzed using a focused beam current of 15 nA. Counting times were 20 s on the peak and 10 s on the background. Glasses and quenched melts were analyzed with a ~10  $\mu$ m defocused beam and counting times were 5 s on the peak and 2.5 s on the background to minimize loss of volatile elements. Matrixmatched minerals were used as reference materials. Additionally, a set of well-characterized secondary standards was measured together with the unknowns to monitor external precision and accuracy (Klemme et al., 2008).

Seventeen trace elements were analyzed using a LA-ICP-MS system at the University of Münster that consisted of an Element 2 (ThermoFisher) SF-ICP-MS connected to a 193 nm ArF excimer laser ablation system (UP193HE, New Wave Research). Laser repetition

Table 2		
Experimental	run	conditions.

Р	T/°C	Run duration/h	Starting material	Phases present
1 atm	1100	28	Mix 2	ol, pv, melt
1 atm	1300	66	Mix 2	ol, pv, melt
1 atm	1200	18	Mix 2	ol, pv, tnt, melt
1 atm	1300	18	Mix 2	ol, pv, melt
1 atm	1350	18	Mix 1	pv, melt
1 atm	1450	18	Mix 1	ol, melt
1.5 GPa	1200	8	MixBeyK1	pv, ol, melt
1.5 GPa	1300	5	SKHDCMB2	pv, melt
	1 atm 1 atm 1 atm 1 atm 1 atm 1 atm 1.5 GPa	1 atm   1100     1 atm   1300     1 atm   1200     1 atm   1300     1 atm   1300     1 atm   1350     1 atm   1450     1.5 GPa   1200	1 atm 1100 28   1 atm 1300 66   1 atm 1200 18   1 atm 1300 18   1 atm 1350 18   1 atm 1450 18   1 atm 1450 18   1 atm 1450 18	1 atm   1100   28   Mix 2     1 atm   1300   66   Mix 2     1 atm   1200   18   Mix 2     1 atm   1300   18   Mix 2     1 atm   1300   18   Mix 1     1 atm   1350   18   Mix 1     1 atm   1450   18   Mix 1     1.5 GPa   1200   8   MixBeyK1

ol = olivine, pv = perovskite, tnt = titanite (sphene), melt = quenched melt. Note that the experiments MCB1\_1 and RPER2\_3 were not analyzed for trace elements as they contained only very small (MCB1\_1) or no perovskite crystals.

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