



Contents lists available at ScienceDirect

Physics Letters A

www.elsevier.com/locate/pla



The bursts of high energy events observed by the telescope array surface detector

R.U. Abbasi^a, M. Abe^b, T. Abu-Zayyad^a, M. Allen^a, R. Anderson^a, R. Azuma^c, E. Barcikowski^a, J.W. Belz^a, D.R. Bergman^a, S.A. Blake^a, R. Cady^a, B.G. Cheon^e, J. Chiba^f, M. Chikawa^g, T. Fujiiⁱ, M. Fukushima^{i,j}, T. Goto^k, W. Hanlon^a, Y. Hayashi^k, N. Hayashida^l, K. Hibino^l, K. Honda^m, D. Ikedaⁱ, N. Inoue^b, T. Ishii^m, R. Ishimori^c, H. Itoⁿ, D. Ivanov^a, C.C.H. Jui^a, K. Kadota^o, F. Kakimoto^c, O. Kalashev^p, K. Kasahara^q, H. Kawai^r, S. Kawakami^k, S. Kawana^b, K. Kawataⁱ, E. Kidoⁱ, H.B. Kim^e, J.H. Kim^a, J.H. Kim^s, S. Kishigami^k, S. Kitamura^c, Y. Kitamura^c, V. Kuzmin^p, Y.J. Kwon^h, J. Lan^a, J.P. Lundquist^a, K. Machida^m, K. Martens^j, T. Matsuda^t, T. Matsuyama^k, J.N. Matthews^a, M. Minamino^k, K. Mukai^m, I. Myers^a, K. Nagasawa^b, S. Nagatakiⁿ, T. Nakamura^u, T. Nonakaⁱ, A. Nozato^g, S. Ogio^k, J. Ogura^c, M. Ohnishiⁱ, H. Ohokaⁱ, K. Okiⁱ, T. Okuda^{v,*}, M. Ono^w, R. Onogi^k, A. Oshima^x, S. Ozawa^q, I.H. Park^y, M.S. Pshirkov^{z,p}, D.C. Rodriguez^a, G. Rubtsov^p, D. Ryu^s, H. Sagawaⁱ, K. Saitoⁱ, Y. Saito^{ae}, N. Sakakiⁱ, N. Sakurai^k, A.L. Sampson^a, L.M. Scott^{aa}, K. Sekinoⁱ, P.D. Shah^a, F. Shibata^m, T. Shibataⁱ, H. Shimodairaⁱ, B.K. Shin^k, H.S. Shinⁱ, J.D. Smith^a, P. Sokolsky^a, R.W. Springer^a, B.T. Stokes^a, S.R. Stratton^{a,aa}, T.A. Stroman^a, T. Suzawa^b, M. Takamura^f, M. Takedaⁱ, R. Takeishiⁱ, A. Taketa^{ab}, M. Takitaⁱ, Y. Tameda^l, H. Tanaka^k, K. Tanaka^{ac}, M. Tanaka^t, S.B. Thomas^a, G.B. Thomson^a, P. Tinyakov^{p,ad}, I. Tkachev^p, H. Tokuno^c, T. Tomida^{ae}, S. Troitsky^p, Y. Tsunesada^k, K. Tsutsumi^c, Y. Uchihori^{af}, S. Udo^l, F. Urban^{ad}, G. Vasiloff^a, T. Wong^a, R. Yamane^k, H. Yamaoka^t, K. Yamazaki^{ab}, J. Yang^d, K. Yashiro^f, Y. Yoneda^k, S. Yoshida^r, H. Yoshii^{ag}, R. Zollinger^a, Z. Zundel^a

^a High Energy Astrophysics Institute and Department of Physics and Astronomy, University of Utah, Salt Lake City, UT, USA

^b The Graduate School of Science and Engineering, Saitama University, Saitama, Saitama, Japan

^c Graduate School of Science and Engineering, Tokyo Institute of Technology, Meguro, Tokyo, Japan

^d Department of Physics and Institute for the Early Universe, Ewha Womans University, Seodaemun-gu, Seoul, Republic of Korea

^e Department of Physics and The Research Institute of Natural Science, Hanyang University, Seongdong-gu, Seoul, Republic of Korea

^f Department of Physics, Tokyo University of Science, Noda, Chiba, Japan

^g Department of Physics, Kinki University, Higashi Osaka, Osaka, Japan

^h Department of Physics, Yonsei University, Seodaemun-gu, Seoul, Republic of Korea

ⁱ Institute for Cosmic Ray Research, University of Tokyo, Kashiwa, Chiba, Japan

^j Kavli Institute for the Physics and Mathematics of the Universe (WPI), Todai Institutes for Advanced Study, The University of Tokyo, Kashiwa, Chiba, Japan

^k Graduate School of Science, Osaka City University, Osaka, Osaka, Japan

^l Faculty of Engineering, Kanagawa University, Yokohama, Kanagawa, Japan

^m Interdisciplinary Graduate School of Medicine and Engineering, University of Yamanashi, Kofu, Yamanashi, Japan

ⁿ Astrophysical Big Bang Laboratory, RIKEN, Wako, Saitama, Japan

^o Department of Physics, Tokyo City University, Setagaya-ku, Tokyo, Japan

^p National Nuclear Research University, Moscow Engineering Physics Institute, Moscow, Russia

^q Advanced Research Institute for Science and Engineering, Waseda University, Shinjuku-ku, Tokyo, Japan

^r Department of Physics, Chiba University, Chiba, Chiba, Japan

^s Department of Physics, School of Natural Sciences, Ulsan National Institute of Science and Technology, UNIST-gil, Ulsan, Republic of Korea

^t Institute of Particle and Nuclear Studies, KEK, Tsukuba, Ibaraki, Japan

^u Faculty of Science, Kochi University, Kochi, Kochi, Japan

^v Department of Physical Sciences, Ritsumeikan University, Kusatsu, Shiga, Japan

^w Department of Physics, Kyushu University, Fukuoka, Fukuoka, Japan

* Corresponding author.

E-mail address: okuda@icrr.u-tokyo.ac.jp (T. Okuda).

^x Engineering Science Laboratory, Chubu University, Kasugai, Aichi, Japan^y Department of Physics, Sungkyunkwan University, Jang-an-gu, Suwon, Republic of Korea^z Sternberg Astronomical Institute, Moscow M.V. Lomonosov State University, Moscow, Russia^{aa} Department of Physics and Astronomy, Rutgers University – The State University of New Jersey, Piscataway, NJ, USA^{ab} Earthquake Research Institute, University of Tokyo, Bunkyo-ku, Tokyo, Japan^{ac} Graduate School of Information Sciences, Hiroshima City University, Hiroshima, Hiroshima, Japan^{ad} Service de Physique Théorique, Université Libre de Bruxelles, Brussels, Belgium^{ae} Department of Computer Science and Engineering, Shinshu University, Nagano, Nagano, Japan^{af} National Institute of Radiological Science, Chiba, Chiba, Japan^{ag} Department of Physics, Ehime University, Matsuyama, Ehime, Japan

ARTICLE INFO

Article history:

Received 12 January 2017

Received in revised form 12 June 2017

Accepted 13 June 2017

Available online xxx

Communicated by P.R. Holland

Keywords:

High energy radiation

Lightning

Terrestrial gamma-ray flash

ABSTRACT

The Telescope Array (TA) experiment is designed to detect air showers induced by ultra high energy cosmic rays. The TA ground Surface particle Detector (TASD) observed several short-time bursts of air shower like events. These bursts are not likely due to chance coincidence between single shower events. The expectation of chance coincidence is less than 10^{-4} for five-year's observation. We checked the correlation between these bursts of events and lightning data, and found evidence for correlations in timing and position. Some features of the burst events are similar to those of a normal cosmic ray air shower, and some are not. On this paper, we report the observed bursts of air shower like events and their correlation with lightning.

© 2017 Elsevier B.V. All rights reserved.

1. Introduction

There have been reports about the observation of energetic radiation from thunderclouds. Some of these reports discuss the increasing rates of radiation at the ground in the presence of thunderclouds. Some of them discuss bursts of energetic radiation observed from space, known as terrestrial gamma-ray flashes (TGF) [1,2], also from aircraft [3]. They are believed to be associated with upward lightning flashes at the tops of thunderclouds. The observation of TGF on the ground was reported by Dwyer, et al. [4] Other observations of energetic radiation have been associated with individual lightning processes in the flash. The possible mechanism for some of these kinds of energetic radiation from thunderclouds is Relativistic Runaway Electron Avalanches (RREAs) or due to strong electric field on the streamer tip summarized by Dwyer, et al. [5]

Lightning is classified by the discharge region, intracloud lightning (IC), cloud to ground lightning (CG) and cloud to cloud lightning. The natural lightning flash consists of several processes, known as the stepped leader, return stroke, dart leader and subsequent return stroke. The leader direction may be up or down and of positive or negative polarity, hence there are four types of lightning for cloud to ground lightning. However, except for lightning strikes on tall objects, most lightning starts with negative charges moving downward.

Moore, et al. [6] reported the observation of the energetic radiation from stepped leaders, using NaI as a radiation detector. Dwyer, et al. [7,8] reported the observation of the energetic radiation from dart leaders, also using NaI as a radiation detector, for rocket triggered lightning. Dwyer, et al. [9] reported the observation of energetic radiation from stepped leaders, and Dwyer, et al. [10] reported association with the return stroke, also using NaI as a radiation detector, for natural cloud to ground lightning.

On the other hand, the stepped leader, the beginning of lightning, cannot start only by the electric fields in the usual thunderstorm. Therefore, it has been hypothesized that cosmic ray air showers play a role in triggering lightning by ionizing the atmosphere. In support of this, there have been several reports of energetic radiation observed with lightning.

Gurevich, et al. [11,12] reported the coincidence of air showers with lightning, using NaI and gas-counters as radiation detectors. Chilingarian, et al. [13] also reported the coincidence of air show-

ers with lightning, using plastic scintillators. Gurevich, et al. [14] presented the results of radio emission measurements and discussed as follows. If cosmic ray air showers stimulate the electron avalanche in lightning, their observed radio event rates seem inconsistent with the flux of cosmic rays with energies estimated from the observed radio amplitudes of electron avalanche. The cosmic ray energies estimated from the avalanche amplitudes are five to six orders of magnitude higher than those estimated from the rates. Hydrometeors are introduced in an effort to overcome this inconsistency [14].

2. Observed burst events

The Telescope Array (TA) experiment, located in midwest Utah, USA(39.3N, 112.9W, Alt 1382 m), consists of two types of detector (Fig. 1). Both methods observe the high energy phenomenon known as an "air shower", which is generated by an ultra high energy cosmic ray. One instrument is atmospheric fluorescence telescope detectors and the other is ground surface particle detectors [15]. In contrast to atmospheric fluorescence which is observable during moonless nights, the TA Surface Detector (TASD) runs all day throughout the year. TASD consists of 507 plastic scintillation counters. The particle detecting part of an individual TASD counter is shown in Fig. 2. The energy deposition on this counter for a vertical muon is 2.043 MeV and the equivalent energy for trigger threshold is 0.7 MeV. The counters are deployed as a square grid array with 1.2 km spacing, and covers 680 km² altogether. When three adjacent detectors detect a signal, each of which corresponds to more than three particle equivalent in 8 μ s, TASD judges that their signals are from an air shower, causing signal waveforms to be digitized from all detectors within $\pm 32 \mu$ s of the trigger time [15]. The TASD is designed to detect all air showers from cosmic rays whose primary energy is more than 10^{19} eV. The average rate of shower trigger is 7.5 mHz and the expected cosmic ray energy at the threshold level is $10^{17.5}$ eV [15].

The particle detectors used in the preceding studies on introduction were mainly NaI, sometimes coupled with gas counters or plastic scintillators. The TASD uses plastic scintillator, with an area approximately 300 times larger than the NaI detectors but lacking the ability to measure the energy of individual particles. The TASD scintillator responds about 10 times faster than NaI. Whereas

Download English Version:

<https://daneshyari.com/en/article/5496299>

Download Persian Version:

<https://daneshyari.com/article/5496299>

[Daneshyari.com](https://daneshyari.com)