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On blow-up criteria for a new Hall-MHD system

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ABSTRACT

This paper investigates a new Hall-MHD system in \mathbb{R}^3 . Besides local well-posedness for strong solutions and global existences for weak solutions, some blow-up criteria are established.

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1. Introduction

In this paper, we consider the following Hall-MHD system [1–4]:

$$\operatorname{div} u = \operatorname{div} b = 0, \tag{1.1}$$

$$\partial_t u + u \cdot \nabla u + \nabla \left(\pi + \frac{1}{2}|b|^2\right) - \Delta u = b \cdot \nabla b, \tag{1.2}$$

$$\partial_t b - \left(\frac{\delta_e}{L_0}\right)^2 \partial_t \Delta b + u \cdot \nabla b - b \cdot \nabla u - \Delta b = \frac{\delta_i}{L_0} \frac{1}{\rho} \operatorname{rot} \left(b \times \operatorname{rot} b\right) - \left(\frac{\delta_e}{L_0}\right)^2 \frac{1}{\rho} \operatorname{rot} \left((u \cdot \nabla) \operatorname{rot} b\right), \tag{1.3}$$

$$(u, b)(\cdot, 0) = (u_0, b_0)(\cdot)$$
 in \mathbb{R}^3 .

Here *u* is the fluid velocity field, π is the pressure and *b* is the magnetic field. L_0 is the normalizing length limit, δ_e and δ_i denote electron and ion inertia respectively, and ρ is the fluid density. For simplicity, we will take $\delta_e = \delta_i = L_0 = \rho = 1$.

The applications of the Hall-MHD system cover a very wide range of physical objects, for example, magnetic reconnection in space plasmas, star formation, neutron stars, and geo-dynamo.

When $\delta_{\ell} = 0$, the system (1.1)–(1.3) reduces to the incompressible Hall-MHD equations, which has been received many studies [5-14]. The paper [5] gave a derivation of (1.1)-(1.3) from a two-fluid Euler-Maxwell system. Chae et al. [8] proved the local existence of smooth solutions. Chae and Lee [6] and Fan and Ozawa [11] proved some regularity criteria. For the compressible and density-dependent Hall-MHD, we refer to [15] and [16].

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When the last term in (1.3) is neglected, the system (1.1)-(1.3) reduced to the two-fluid MHD equation. Chen and Miao [17] proved the local existence of strong solutions and some regularity criteria.

The aim of this paper is to investigate well-posedness and blow-up criteria for (1.1)-(1.4). First, we have the following local well-posedness theorem for the strong solution.

Theorem 1.1. Let $u_0 \in H^s$, $b_0 \in H^{s+1}$ with $s \ge 3$ and div $u_0 = \text{div } b_0 = 0$. Then there exist T > 0 and a unique strong solution (u, b) to the problem (1.1)-(1.4) satisfying

$$\mu \in L^{\infty}(0,T;H^{s}) \cap L^{2}(0,T;H^{s+1}), \ b \in L^{\infty}(0,T;H^{s+1}).$$
(1.5)

Since the proof of Theorem 1.1 is very similar to that in [17], we omit the details here. At the same time, it is easy to establish a global existence for weak solutions similarly to that for the standard MHD equations.

Theorem 1.2. Let $u_0 \in L^2$, $b_0 \in H^1$ and div $u_0 = \text{div } b_0 = 0$. Then the problem has a weak solution u, b satisfying energy inequality for any T > 0.

Note that the last term in (1.3) can be rewritten as $-\sum_{i=1}^{3} \operatorname{rot} (\partial_i (u_i \operatorname{rot} b))$, but we omit the detailed proof here. The following blow-up criteria are the main results and will be proved in Section 2.

Theorem 1.3. Let $u_0 \in H^s$, $b_0 \in H^{s+1}$ with $s \ge 3$ and div $u_0 = \text{div } b_0 = 0$. Let (u, b) be the unique solution constructed in Theorem 1.1. If one of the following two conditions

(i)
$$\int_{0}^{1} (\|\nabla u(t)\|_{L^{\infty}} + \|b(t)\|_{BMO}^{2}) dt < \infty,$$
 (1.6)

(ii)
$$\int_{0}^{T} (\|\nabla u(t)\|_{\dot{B}^{0}_{\infty,\infty}} + \|\nabla b(t)\|_{\dot{B}^{0}_{\infty,\infty}}) dt < \infty,$$
(1.7)

holds for some $0 < T < \infty$, then the solution (u, b) can be extended beyond T > 0.

Here $\dot{B}_{\infty,\infty}^0$ denotes the homogeneous Besov space and *BMO* is the space of bounded mean oscillation.

Definition 1.1. (see [18]). Let $\{\phi_j\}_{j\in\mathbb{Z}}$ be the Littlewood–Paley dyadic decomposition of unity that satisfies $\hat{\phi} \in C_0^{\infty}(B_2 \setminus B_{1/2})$, $\hat{\phi}_j(\xi) = \hat{\phi}(2^{-j}\xi)$ and $\sum_{j\in\mathbb{Z}} \hat{\phi}_j(\xi) = 1$ for any $\xi \neq 0$, where $\hat{\phi}$ is the Fourier transform of ϕ and B_r is the ball with radius r centered at the origin. The homogeneous Besov space is defined as

$$\dot{B}^{s}_{p,q} := \left\{ f \in \mathcal{S}' / \mathcal{P} : \|f\|_{\dot{B}^{s}_{p,q}} < \infty \right\}$$

with the norm

$$\|f\|_{\dot{B}^s_{p,q}} \coloneqq egin{cases} \left\{ igl(\sum_{j\in\mathbb{Z}}\|2^{js}\Delta_jf\|_{L^p}^qigr)^{rac{1}{q}}, & 1\leq q<\infty \ \sup_{j\in\mathbb{Z}}\|2^{js}\Delta_jf\|_{L^p}, & q=\infty, \end{cases}
ight.$$

for all $s \in \mathbb{R}$ and $1 \le p \le \infty$ and $\Delta_i f := \phi_i * f$, where S' is the space of tempered distributions and \mathcal{P} is the space of polynomials.

We will also use the following logarithmic Sobolev inequalities [19,20]:

$$\|\nabla u\|_{L^{\infty}} \le C(1 + \|\nabla u\|_{\dot{B}^{0}_{\infty,\infty}} \log (e + \|u\|_{H^{s}})), \tag{1.8}$$

$$\|b\|_{L^{\infty}} \le C(1+\|b\|_{BMO} \log^{1/2} (e+\|b\|_{H^2})), \tag{1.9}$$

with $s > \frac{5}{2}$.

2. Proof of Theorem 1.3

This section is devoted to the proof of Theorem 1.3. We only need to establish a priori estimates. First, we can do the following energy estimates.

Testing (1.2) by *u* and using (1.1), we see that

$$\frac{1}{2}\frac{d}{dt}\int |u|^2 dx + \int |\nabla u|^2 dx = \int (b \cdot \nabla)b \cdot u dx.$$
(2.1)

Testing (1.3) by *b* and using (1.1), we infer that

$$\frac{1}{2}\frac{d}{dt}\int (|b|^2 + |\nabla b|^2)dx + \int |\nabla b|^2dx = \int (b \cdot \nabla)u \cdot bdx.$$
(2.2)

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