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Alexander V. Rodionov

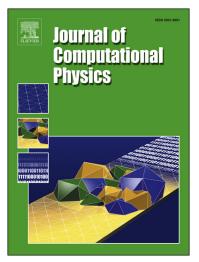
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## ACCEPTED MANUSCRIPT

#### Artificial viscosity to cure the carbuncle phenomenon: the three-dimensional case

### Alexander V. Rodionov

Russian Federal Nuclear Center – All-Russian Scientific Research Institute of Experimental Physics (RFNC-VNIIEF), Sarov, Nizhny Novgorod region, 607190, Russia

E-mail address: avrodionov@rambler.ru

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#### 1. Introduction

The *carbuncle phenomenon* (also known as the *shock instability*) has remained a serious computational challenge since it was first noticed and described [1, 2]. In [3] the author presented a summary on this subject and proposed a new technique for curing the problem. Its idea is to introduce some dissipation in the form of right-hand sides of the Navier-Stokes equations into the basic method of solving Euler equations; in so doing, the molecular viscosity coefficient is replaced by the artificial viscosity coefficient. The new cure for the carbuncle flaw was tested and tuned for the case of using first-order schemes in two-dimensional simulations. Its efficiency was demonstrated on several well-known test problems. In this paper we extend the technique of [3] to the case of three-dimensional simulations.

## 2. Governing equations and space discretization

<u>The Euler equations.</u> The governing equations for the compressible gas dynamics, written in Cartesian coordinates *xyz*, are

$$\frac{\partial \mathbf{U}}{\partial t} + \frac{\partial \mathbf{F}_{x}}{\partial x} + \frac{\partial \mathbf{F}_{y}}{\partial y} + \frac{\partial \mathbf{F}_{z}}{\partial z} = \text{RHS}, \qquad \mathbf{U} = \begin{bmatrix} \rho \\ \rho u_{x} \\ \rho u_{y} \\ \rho u_{z} \\ \rho h_{0} - P \end{bmatrix}, \qquad \mathbf{F}_{m} = \begin{bmatrix} \rho u_{m} \\ \rho u_{x} u_{m} + \delta_{mx} P \\ \rho u_{y} u_{m} + \delta_{my} P \\ \rho u_{z} u_{m} + \delta_{mz} P \\ \rho h_{0} u_{m} \end{bmatrix}, \qquad m = x, y, z, \qquad (1)$$

where RHS = 0 for the Euler equations,  $\delta_{mn}$  is the Kronecker delta,  $\mathbf{u} = (u_x, u_y, u_z)$  are the velocity vector and its components,  $\rho$  is the density, P is the pressure and  $h_0$  is the specific total enthalpy, which for a polytropic gas is

$$h_0 = \frac{1}{2} \left( u_x^2 + u_y^2 + u_z^2 \right) + h, \quad h = \frac{\gamma P}{(\gamma - 1)\rho},$$

with *h* denoting the specific enthalpy and  $\gamma$  denoting the ratio of specific heats.

<u>The space discretization</u>. We apply the finite volume approach to solve (1) numerically. In so doing we consider a sufficiently smooth structured grid and introduce curvilinear coordinates  $\xi \eta \zeta$  that transform the grid in physical space *xyz* to a rectangular grid in computational space with the grid spacing  $\Delta \xi = \Delta \eta = \Delta \zeta = 1$ . The following geometric parameters of the grid will take the place of metrics (hereinafter the grid indices *i*, *j* and *k* correspond to the coordinates  $\xi$ ,  $\eta$  and  $\zeta$ ):  $V_{ij,k}$ , the cell volume;  $(S_{\xi})_{i+\frac{1}{2},j,k}$ ,  $(S_{\eta})_{i,j+\frac{1}{2},k}$  and  $(S_{\zeta})_{i,j,k+\frac{1}{2}}$ , the area vectors of the cell faces in three grid directions (each area vector points in the direction of increase in the associated coordinate).

Assume that the solution at a given time level *n* is defined, meaning that for the time  $t = t^n$  we know the average values  $\mathbf{Q}_{i,j,k}^n \equiv (u_x, u_y, u_z, \rho, P)_{i,j,k}^n$  within each cell. In order to update the solution

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