

Contents lists available at ScienceDirect

Materials Science in Semiconductor Processing

journal homepage: www.elsevier.com/locate/mssp



Check fo

In situ growth of CuSbS₂ thin films by reactive co-sputtering for solar cells

Liangliang Kang^a, Lianbo Zhao^a, Liangxing Jiang^{a,c,*}, Chang Yan^c, Kaiwen Sun^c, Boon K. Ng^b, Chunhui Gao^a, Fangyang Liu^{a,c,*}



^b Faculty of EsTEM, University of Canberra, Bruce, ACT 2601, Australia

^c Australian Centre for Advanced Photovoltaics, School of Photovoltaic and Renewable Energy Engineering, University of New South Wales, Sydney 2052, Australia

ARTICLE INFO	ABSTRACT				
Keywords:	-CuSbS ₂ thin films were in situ grown by reactive co-sputtering and the effects of the growth temperature on film				
Reactive sputtering	composition, structure and morphology were investigated. It is demonstrated that orthorhombic chalcostibite				
In situ growth CuSbS ₂ Thin films Photovoltaics	$CuSbS_2$ thin films with uniform morphology, pure phase and grain size over 2 µm can be obtained for growth temperature of 300 °C, while higher growth temperature results in the formation of Cu_3SbS_4 phase and lower				
	growth temperature leads to Sb_2S_3 secondary phase. The grown CuSbS_2 film shows an optical absorption coefficient of higher than 10^4 cm ⁻¹ , an optical band gap of 1.52 eV and p-type conductivity. Solar cell devices				
	with configuration of glass/Mo/CuSbS2/CdS/i-ZnO/ITO/Ag were fabricated and yield power conversion effi-				
	ciency of 0.52%. The incompatible interfaces including absorber/back contact (the absence of beneficial MoS_2				
	interface layer) and absorber/buffer layer (unfavorable "cliff"-like conduction band offset) interfaces have been				

considered as the key factor limiting efficiency.

1. Introduction

In recent years, CdTe and CuInGaSe2 (CIGSe) have been considered as promising absorber materials for thin film photovoltaics (PV) with high conversion efficiency of above 20% [1,2]. However, the scarcity of Te, In, and Ga are major problems limiting the widespread utilization of these devices. Hence, the development of low cost, low toxicity, earth abundant and long-term stability photovoltaic materials is critical for thin film solar cell manufacture. Recently, ternary I-V-VI2 chalcogenide CuSbS₂, which has high optical absorption coefficient between 10⁴ and 10^{6} cm⁻¹ in the visible region [3,4] and a direct band gap of 1.4–1.6 eV [4-6], has aroused interest for application as a promising innovative inorganic absorber material for sustainable and propagable PV due to its non-toxic, low cost, long-term stability and earth-abundant constituents [3]. Furthermore, CuSbS₂ has a relatively low melting point of 551 °C [7], indicating that it is suited to crystallization at lower temperatures. Therefore, the exploration of CuSbS₂ for photovoltaic application was of great importance based on all these characteristics. Power conversion efficiencies ~ 3% level have been achieved for CuSbS₂ thin film solar cells with glass/Mo/CuSbS₂/CdS/i-ZnO/ITO/Ag device configuration [8,9]. Besides, 4.70% efficiency has been reach for selenide CuSbSe₂ thin film devices, which is the best performance for I-V-VI₂ chalcogenide based solar cells up to now [10].

CuSbS₂ thin films can be fabricated by a series of deposition

techniques such as thermal evaporation [11-15], RF sputtering [16-18], atmospheric plasma chemical vapor deposition (APCVD) [19], atomic layer deposition (ALD) [20], sulfurization of metal precursor [21–24], spray pyrolysis deposition (SPD) [25–28], hybrid inks [8,29], hot-injection [30-32], hydrothermal [33], solvothermal [34,35], chemical bath deposition (CBD) [36-38], electrodeposition [9,39-41], Spin-coating [42] and drop-casting [43], etc. Deposition of high-quality CuSbS₂ thin films is enormously challenging could be attributed to (i) a large number of ternary copper antimony sulfide phases (i.e., Cu₃SbS₃ [16], Cu₃SbS₄ [44], Cu₁₂Sb₄S₁₃ [39]) can be formed depending on the process conditions or stoichiometry used during fabrication, and/or (ii) other binary impurities can be synthesized or segregated, such as Cu_xS [30] and Sb_2S_3 [45]. The aforementioned deposition techniques, which all involve the precursors preparation followed by a high temperature sulfurization, may easily introduce impure phase into the film and increase the complexity for large scale applications. It is necessary that a large area deposition process and direct "one step" should be applied in the field of large scale production of thin film PV. Reactive co-sputtering deposition techniques, as a suitable technique for this task, can not only simply provide control on deposition rate, thickness and film composition by just adjusting the deposition time and ratio of the powers applied to the metal targets, but also applicable for large area, successive and multi-component film deposition with relatively low cost [46]. In addition, the phase of material can be regulated in real

https://doi.org/10.1016/j.mssp.2018.05.004 Received 26 February 2018; Received in revised form 24 April 2018; Accepted 3 May 2018 1369-8001/ © 2018 Published by Elsevier Ltd.

^{*} Corresponding authors at: School of Metallurgy and Environment, Central South University, Changsha 410083, China. E-mail addresses: lxjiang_csu@163.com (L. Jiang), liufangyang@csu.edu.cn (F. Liu).

time to induce the growth of a small amount of impurities, defects and secondary phases in the film through this method [47].

In this work, we present a Reactive co-sputtering deposition techniques for the in situ growth of $CuSbS_2$ thin films onto Mo-coated glass substrates for solar cells. The effects of the growth temperature on film composition, structure and morphology were investigated. Solar cell device with configuration of glass/Mo/CuSbS_2/CdS/i-ZnO/ITO/Ag was also fabricated and yielded a conversion efficiency of 0.52%.

2. Experimental

CuSbS₂ films were deposited onto Mo-coated glass substrates (40 mm × 40 mm) by reactive co-sputtering with two metal targets of Cu (purity: 99.99%) and Sb (purity: 99.99%) in the mixed Ar/H₂S working gas at different growth temperatures (200–350 °C). The metal targets were 5 mm in thickness and 60 mm in diameter. Before sputtering, the chamber was pumped down to lower than 1×10^{-4} Pa. The flow rate of argon (purity: 99.999%) and hydrogen sulfide (purity: 99.9%) were 30 and 10 sccm, respectively. The working pressure is set at 0.5 Pa. The sputtering source was operated using metallic Cu and Sb targets, by two independent RF power of 120 W and 80 W, respectively. The sputtering time was 15 min. The CuSbS₂ solar cells were completed using the CBD 50 nm CdS buffer layer, the RF-sputtering 50 nm intrinsic ZnO (i-ZnO) and the DC-sputtering 300 nm ITO window layer. Each device had a total area of approximately 0.45 cm² defined by mechanical scribing.

The composition of films was investigated by X-ray fluorescence (XRF, Shimadzu XRF-1800). The morphology of the thin films were studied by scanning electron microscopy (SEM, FEI Quanta-200). The phase structure was measured by both X-ray diffraction (XRD, Rigaku3014) patterns and Raman spectra (Raman, LabRAM HR800, 520 nm). The elements state of the CuSbS₂ films were analyzed by X-ray Photoelectron Spectroscopy (XPS, KAlpha). Optical properties of the films were measured using an UV–Vis–NIR spectrophotometer (Hitachi U-4100). The electrical properties of samples were measured using Hall effects measurement (HMS-3000/0.55T). Current density-voltage (*J-V*) characterization for PV structures were performed using an AM 1.5 solar simulator (NEWPORT, 100 mW/cm²).

3. Results and discussion

The elemental composition of the grown thin films at different growth temperature was determined by XRF measurements and shown in Table 1. It was observed that the composition of the deposited films is sensitive to the growth temperature. The Cu/Sb radios of the films increases 0.48–2.70 as the growth temperature increased from 200 °C to 350 °C, which may be related to the high volatility of Sb₂S₃ phase [3,11]. For growth temperature of 200 °C and 250 °C, atomic ratio Cu/Sb is far lower than 1, which is close to 1 and shows slightly Cu-poor composition at growth temperature of 300 °C. Cu/Sb rises sharply to highly Cu-rich when the growth temperature reaches 350 °C, suggesting a significant loss in Sb. The S/(Cu+Sb) radios were observed to be reduced with increasing in growth temperature from 200 °C to 300 °C, and then remain at 1.1–1.2, showing a slightly S-rich composition.

Table 1

Elemental composition of the $CuSbS_2$ films prepared at different growth temperature.

S.No.	growth temperature	Atomic percent (at%)			Ratio	
		Cu	Sb	S	Cu/Sb	S/(Cu+Sb)
1	200 °C	13.34	28.06	58.60	0.48	1.41
2	250 °C	15.90	28.94	55.15	0.55	1.22
3	300 °C	23.05	24.25	52.70	0.95	1.11
4	350 °C	33.80	12.52	53.68	2.70	1.15



Fig. 1. XRD diffraction patterns for $CuSbS_2$ thin films deposited at growth temperature of 200, 250, 300 and 350 °C.

According to previous report, Cu-poor and S-rich composition is required for high efficiency I-V-VI₂ chalcogenide based solar cells [3], the optimal growth temperature is 300 °C.

Fig. 1 shows XRD of the grown thin films at different growth temperature. It is obvious that the phase formation was influenced strongly by the growth temperature. No diffraction peaks of CuSbS₂ can be observed in the thin films grown at 200 °C except for the peaks from substrate, indicating an amorphous state of the film. For growth temperature of 250 °C, the samples shows orthorhombic chalcostibite CuSbS₂ (JCPDS No 44-1417) phase with segregated Sb₂S₃ [11,48] (JCPDS No 42-1393), which explains the atomic ratio of Cu/Sb far lower than 1. For growth temperature of 300 °C the sample shows only orthorhombic CuSbS₂ phase without observable secondary phases. Two intense and sharp peaks at about 28.4° and 29.9° from (111) and (301) reflections can be observed, revealing good crystallinity. For growth temperature of 350 °C, the samples showed main phase of the famatinite Cu₃SbS₄ [34,43] (JCPDS No 35–0581), which leads to the high Cu/ Sb value of 2.7 as show in Table 1. Therefore, 300 °C again can be considered as the most suitable growth temperature.

Fig. 2 shows SEM images of surface morphologies of the CuSbS₂ thin films prepared at different growth temperatures. The films grown at 200-300 °C shows compact and uniform surface (Fig. 2(a-c)). And it can be seen that the grain boundaries become clear and grains size become larger gradually with the increasing in growth temperature, suggesting an improvement in crystallinity. Particularly at growth temperature of 300 °C as shown in Fig. 2(c), the film consisted of dense large grain with size of larger than $2\,\mu m$, which is beneficial to high efficiency of the thin film solar cells according to the experience in CIGSe photovoltaics [49]. As the growth temperature was further increases to 350 °C, the film surface microstructure (Fig. 2(d)) has changed dramatically, showing cluster-like morphology consisting of belted-particles. Combined with XRF and XRD analysis, this can be related to phase transformation of the films from CuSbS₂ to Cu₃SbS₄. Based on the above composition, structure and morphology characterizations, one can clearly find that growth temperature of 300 °C can yield orthorhombic CuSbS₂ thin films with slightly Cu-poor and S-rich composition, compact and uniform morphology, and good crystallinity, which was employed for further characterizations and device fabrication.

Fig. 3(a) displays the cross-section of the films prepared at growth temperature of 300 °C. It is obviously that film has a thickness of 1 μ m and shows large columnar grains but a fine grain layer at the bottom. No voids in the CuSbS₂ absorber or at the absorber/Mo interface can be observed. The interface layer MoS₂, which is usually found in CIGSe and Cu₂ZnSnS₄ (CZTS) cases, also cannot be observed. This suggests the interface reactions of Mo with absorber and/or chalcogen from the reactive annealing, which occur in CIGSe and CZTS cases, is not

Download English Version:

https://daneshyari.com/en/article/7117562

Download Persian Version:

https://daneshyari.com/article/7117562

Daneshyari.com