Accepted Manuscript

Energy Spectrum of Fast Second Order Fermi Accelerators as Sources of Ultra-High-Energy Cosmic Rays

Tobias Winchen, Stijn Buitink

 PII:
 S0927-6505(17)30267-0

 DOI:
 10.1016/j.astropartphys.2018.04.004

 Reference:
 ASTPHY 2290

To appear in: Astroparticle Physics

Received date:	18 September 2017
Revised date:	15 February 2018
Accepted date:	17 April 2018

Please cite this article as: Tobias Winchen, Stijn Buitink, Energy Spectrum of Fast Second Order Fermi Accelerators as Sources of Ultra-High-Energy Cosmic Rays, *Astroparticle Physics* (2018), doi: 10.1016/j.astropartphys.2018.04.004

This is a PDF file of an unedited manuscript that has been accepted for publication. As a service to our customers we are providing this early version of the manuscript. The manuscript will undergo copyediting, typesetting, and review of the resulting proof before it is published in its final form. Please note that during the production process errors may be discovered which could affect the content, and all legal disclaimers that apply to the journal pertain.



Energy Spectrum of Fast Second Order Fermi Accelerators as Sources of Ultra-High-Energy Cosmic Rays

Tobias Winchen^{a,*}, Stijn Buitink^a

^a Vrije Universiteit Brussel, Pleinlaan 2, 1050 Brussels, Belgium

Abstract

Stochastic acceleration of cosmic rays in second order Fermi processes is usually considered too slow to reach ultra-high energies, except in specific cases. In this paper we present the energy spectrum obtained from second order Fermi acceleration in highly turbulent magnetic fields as e.g. found in the outskirts of AGN jets in situations where it can be sufficiently fast to accelerate particles to the highest observed energies. We parametrize the resulting non-power-law spectra and show that these can describe the cosmic ray energy spectrum and mass-composition data at the highest energies if propagation effects are taken into account.

Keywords: High energy cosmic rays, UHECR, Acceleration of particles, Fermi-acceleration, Cosmic ray sources, Spectrum, Hillas' Plot

1. Introduction

Cosmic rays are observed with energies from approximately $10 \,\text{GeV}$ up to energies above $100 \,\text{EeV}$ with a distribution commonly parametrized as a

Preprint submitted to Elsevier

April 17, 2018

^{*}Corresponding author

Email address: tobias.winchen@rwth-aachen.de (Tobias Winchen)

Download English Version:

https://daneshyari.com/en/article/8132665

Download Persian Version:

https://daneshyari.com/article/8132665

Daneshyari.com