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WKB based methods are in common between the alpha-decay and proton-decay models. In this frame work the most important differences between these two decay modes are: the inter-action between the radiated particles and the daughter nucleus and the formation factor of the з exiting particle which in the case of proton-decay is the spectroscopic factor. The spectroscopic factor can be calculated using the relativistic mean field model [37,38] or BCS model [39–43].

An essential part of the WKB calculations which makes them self consistent is the Bohr– Sommerfeld quantization condition [44]. Orientation angle dependent potentials have been widely used in the alpha-decay [45–52], proton decay [24,25,28,29,53], and heavy ion reaction [54–56] studies. In Ref. [49] has been proposed a method which can be used in order to apply the Bohr–Sommerfeld quantization condition to the case of alpha emitters. Ref. [57] has used this method in the case of cluster decay and in this paper we will use the same method in order to cal-culate the half lives of some proton emitters. In this method the normal nuclear potential should be multiplied by an extra angle dependent coefficient which is determined using the quantization condition at each angle. Recently, Santhosh and Sukumaran [53] have applied the Coulomb and proximity potential model for deformed nuclei (CPPMDN) to proton radioactivity by taking into account the deformation of parent and daughter nuclei.

In the first section we summarize the theory of the proton decay based on tunneling model. The numerical results and discussions are presented in the second section and finally the conclusions will be expressed.

2. Model

2.1. Potential

The potential energy that a single proton feels in a spherical nucleus can be estimated using mean field potentials such as the real part of an optical model potential [58,12]:

$V = V_N + V_{Spin-Orbit} + V_{Coul} + V_L$	
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$$V_N = -V_R f(r, R_R, a_R)$$

 $V_R = [54.0 - 0.32E_P + 0.4\frac{Z}{1} + 24.0\frac{(N-Z)}{2}]$ MeV

$$f(r, R, a) = \frac{1}{\left[1 + e^{(r-R)/a}\right]}^{A^3}$$

$$V_{Spin-Orbit} = V_{SO} \overrightarrow{\sigma} \cdot \overrightarrow{L} \lambda_{\pi}^2 \frac{1}{r} \frac{d}{dr} f(r, R_{SO}, a_{SO})$$

$$\overrightarrow{\sigma} \cdot \overrightarrow{L} = L, \quad (j = L + \frac{1}{2})$$
³⁷
₃₈

$$\overrightarrow{\sigma} \cdot \overrightarrow{L} = -(L+1), \quad (j = L - \frac{1}{2} > 0)$$
³⁹
⁴⁰

$$V_{Coul} = (Ze^2 / 8\pi\epsilon_0 R_c) [3 - r^2 / R_c^2], \quad (r \le R_c)$$

$$V_{Coul} = Ze^2 / 4\pi\epsilon_0 r, \quad (r > R_c)$$

$$V_L = L(L+1)\frac{\hbar^2}{2mr^2}$$
(1)
42
43
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in which Z, N and A are the charge, the neutron number and the mass number of the daughter nucleus respectively. The constants have the following values [12]: $R_R = 1.17A^{1/3}$, $a_R = 0.75$,

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