



# Causality and Legendrian linking for higher dimensional spacetimes

Vladimir Chernov

Department of Mathematics, 6188 Kemeny Hall, Dartmouth College, Hanover, NH 03755, USA

## ARTICLE INFO

### Article history:

Received 9 June 2018

Accepted 20 June 2018

### Keywords:

Globally hyperbolic spacetime

Light ray

Contact structure

Legendrian linking

Causality

Legendrian Low conjecture

## ABSTRACT

Let  $(X^{m+1}, g)$  be an  $(m+1)$ -dimensional globally hyperbolic spacetime with Cauchy surface  $M^m$ , and let  $\tilde{M}^m$  be the universal cover of the Cauchy surface. Let  $\mathcal{N}_X$  be the contact manifold of all future directed unparameterized light rays in  $X$  that we identify with the spherical cotangent bundle  $ST^*M$ . Jointly with Stefan Nemirovski we showed that when  $\tilde{M}^m$  is **not** a compact manifold, then two points  $x, y \in X$  are causally related if and only if the Legendrian spheres  $\mathfrak{S}_x, \mathfrak{S}_y$  of all light rays through  $x$  and  $y$  are linked in  $\mathcal{N}_X$ .

In this short note we use the contact Bott–Samelson theorem of Frauenfelder, Labrousse and Schlenk to show that the same statement is true for all  $X$  for which the integral cohomology ring of a closed  $\tilde{M}$  is **not** the one of the CROSS (compact rank one symmetric space).

If  $M$  admits a Riemann metric  $\bar{g}$ , a point  $x$  and a number  $\ell > 0$  such that all unit speed geodesics starting from  $x$  return back to  $x$  in time  $\ell$ , then  $(M, \bar{g})$  is called a  $Y_\ell^X$  manifold. Jointly with Stefan Nemirovski we observed that causality in  $(M \times \mathbb{R}, \bar{g} \oplus -t^2)$  is **not** equivalent to Legendrian linking. Every  $Y_\ell^X$ -Riemann manifold has compact universal cover and its integral cohomology ring is the one of a CROSS. So we conjecture that Legendrian linking is equivalent to causality if and only if one can **not** put a  $Y_\ell^X$  Riemann metric on a Cauchy surface  $M$ .

© 2018 Elsevier B.V. All rights reserved.

All manifolds, maps etc. are assumed to be smooth unless the opposite is explicitly stated, and the word *smooth* means  $C^\infty$ .

## 1. Introduction

Let  $M$  be a not necessarily orientable, connected manifold of dimension  $m \geq 2$  and let  $\pi_M : ST^*M \rightarrow M$  be its spherical cotangent bundle. The manifold  $ST^*M$  carries a canonical co-oriented contact structure. An isotopy  $\{L_t\}_{t \in [0,1]}$  of Legendrian submanifolds in a co-oriented contact manifold is called respectively *positive*, *non-negative* if it can be parameterized in such a way that the tangent vectors of all the trajectories of individual points lie in respectively positive, non-negative tangent half-spaces defined by the contact structure.

For the introduction of basic notions from Lorentz geometry we follow our paper [1].

Let  $(X^{m+1}, g)$  be an  $(m+1)$ -dimensional Lorentz manifold and  $x \in X$ . A nonzero  $\mathbf{v} \in T_x X$  is called *timelike*, *spacelike*, *non-spacelike (causal)* or *null (lightlike)* if  $g(\mathbf{v}, \mathbf{v})$  is respectively negative, positive, non-positive or zero. An piecewise smooth curve is timelike if all of its velocity vectors are timelike. Null and non-spacelike curves are defined similarly. The Lorentz

E-mail address: [Vladimir.Chernov@dartmouth.edu](mailto:Vladimir.Chernov@dartmouth.edu).

manifold  $(X, g)$  has a unique Levi-Civita connection, see for example [2, page 22], so we can talk about timelike and null geodesics. A submanifold  $M \subset X$  is *spacelike* if  $g$  restricted to  $TM$  is a Riemann metric.

All non-spacelike vectors in  $T_x X$  form a cone consisting of two hemicones, and a continuous with respect to  $x \in X$  choice of one of the two hemicones (if it exists) is called the *time orientation* of  $(X, g)$ . The vectors from the chosen hemicones are called *future directed*. A time oriented Lorentz manifold is called a *spacetime* and its points are called *events*.

For  $x$  in a spacetime  $(X, g)$  its *causal future*  $J^+(x) \subset X$  is the set of all  $y \in X$  that can be reached by a future pointing causal curve from  $x$ . The causal past  $J^-(x)$  of the event  $x \in X$  is defined similarly.

Two events  $x, y$  are said to be *causally related* if  $x \in J^+(y)$  or  $y \in J^+(x)$ .

A spacetime is said to be *globally hyperbolic* if  $J^+(x) \cap J^-(y)$  is compact for every  $x, y \in X$  and if it is *causal*, i.e. it has no closed non-spacelike curves. The classical definition of global hyperbolicity requires  $(X, g)$  to be strongly causal rather than just causal, but these two definitions are equivalent, see Bernal and Sanchez [3, Theorem 3.2].

A *Cauchy surface* in  $(X, g)$  is a subset such that every inextendible nonspacelike curve  $\gamma(t)$  intersects it at exactly one value of  $t$ . A classical result is that  $(X, g)$  is globally hyperbolic if and only if it has a Cauchy surface, see [4, pages 211–212]. Geroch [5] proved that every globally hyperbolic  $(X, g)$  is homeomorphic to a product of  $\mathbb{R}$  and a Cauchy surface. Bernal and Sanchez [6, Theorem 1], [7, Theorem 1.1], [8, Theorem 1.2] proved that every globally hyperbolic  $(X^{m+1}, g)$  has a smooth spacelike Cauchy surface  $M^m$  and that moreover for every smooth spacelike Cauchy surface  $M$  there is a diffeomorphism  $h : M \times \mathbb{R} \rightarrow X$  such that

- a:  $h(M \times t)$  is a smooth spacelike Cauchy surface for all  $t$ ,
- b:  $h(x \times \mathbb{R})$  is a future directed timelike curve for all  $x \in M$ , and finally
- c:  $h(M \times 0) = M$  with  $h|_{M \times 0} : M \rightarrow M$  being the identity map.

For a spacetime  $X$  we consider its space of light rays  $\mathfrak{N} = \mathfrak{N}_X$ . By definition, a point  $\gamma \in \mathfrak{N}$  is an equivalence class of inextendible future-directed null geodesics up to an orientation preserving affine reparametrization.

A seminal observation of Penrose and Low [9–11] is that the space  $\mathfrak{N}$  has a canonical structure of a contact manifold (see also [12–15]). A contact form  $\alpha_M$  on  $\mathfrak{N}$  defining that contact structure can be associated to any smooth spacelike Cauchy surface  $M \subset X$ . Namely, consider the map

$$\iota_M : \mathfrak{N}_X \hookrightarrow T^*M$$

taking  $\gamma \in \mathfrak{N}$  represented by a null geodesic  $\gamma \subset X$  to the 1-form on  $M$  at the point  $x = \gamma \cap M$  collinear to  $\langle \dot{\gamma}(x), \cdot \rangle|_M$  and having unit length with respect to the induced Riemann metric on  $M$ . This map identifies  $\mathfrak{N}$  with the unit cosphere bundle  $ST^*M$  of the Riemannian manifold  $M$ . Then

$$\alpha_M := \iota_M^* \lambda_{\text{can}},$$

where  $\lambda_{\text{can}} = \sum p_k dq^k$  is the canonical Liouville 1-form on  $T^*M$ .

**Remark 1.1** (*Bott–Samelson Type Result of Frauenfelder, Labrousse and Schlenk and its Strengthening*). The contact Bott–Samelson [16] type result of Frauenfelder, Labrousse and Schlenk [17, Theorem 1.13] says that if there is a positive Legendrian isotopy of a fiber  $S_x$  of  $ST^*M$  to itself, then the universal cover  $\tilde{M}$  of  $M$  is compact and has the integral cohomology ring of a CROSS. (This result answers our question with Nemirovski [18, Example 8.3] and compactness of  $M$  was first proved in [18, Corollary 8.1].)

Our result with Nemirovski [19, Proposition 4.5] says that if there is a non-constant non-negative Legendrian isotopy of  $S_x$  to itself, then there is a positive Legendrian isotopy of  $S_x$  to itself. (Note that this positive Legendrian isotopy generally is not a perturbation of the non-negative non-constant Legendrian isotopy that was assumed.)

So we can somewhat strengthen [17, Theorem 1.13] to say that if there is a non-constant non-negative Legendrian isotopy of  $S_x$  to itself, then the universal cover  $\tilde{M}$  of  $M$  is compact and has the integral cohomology ring of a CROSS, and in this work we will have to use the strengthened version of the contact Bott–Samelson type theorem.

## 2. Main results

Let  $(X, g)$  be a globally hyperbolic spacetime with Cauchy surface  $M$ . For a point  $x \in X$  we denote by  $\mathfrak{S}_x \subset \mathfrak{N}$  the Legendrian sphere of all (unparameterized, future directed) light rays passing through  $x$ .

For two causally unrelated points  $x, y \in X$  the Legendrian link  $(\mathfrak{S}_x, \mathfrak{S}_y)$  in  $\mathcal{N}$  does not depend on the choice of the causally unrelated points. Under the identification  $\mathcal{N} = ST^*M$  this link is Legendrian isotopic to the link of sphere-fibers over two points of some (and then any) spacelike Cauchy surface  $M$ , see [20, Theorem 8], [1, Lemma 4.3] and [12]. We call a Legendrian link *trivial* if it is isotopic to such a link.

In [1, Theorem A] and [18, Theorem 10.4] we proved the following result. Assume that the universal cover  $\tilde{M}$  of a Cauchy surface of  $M$  of  $X$  is not compact and events  $x, y \in X$  are causally related. Then the Legendrian link  $(\mathfrak{S}_x, \mathfrak{S}_y)$  is nontrivial.

In the case where  $M = \mathbb{R}^3$  this proved the Legendrian Low conjecture of Nataro and Tod [12]. The question to explore relations between causality and linking was motivated by the observations of Low [10, 11] and appeared on Arnold’s problem lists as a problem communicated by Penrose [21, Problem 8], [22, Problem 1998-21].

In this work we prove the following theorem.

Download English Version:

<https://daneshyari.com/en/article/8255305>

Download Persian Version:

<https://daneshyari.com/article/8255305>

[Daneshyari.com](https://daneshyari.com)