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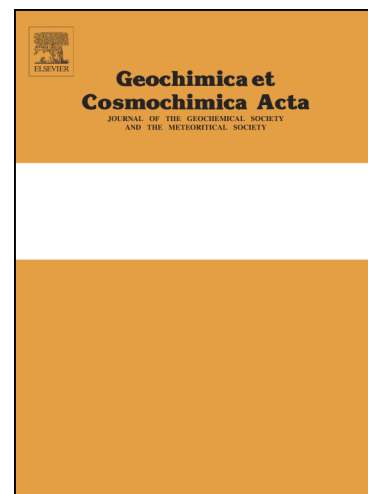
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Chronology of formation of early solar system solids from bulk Mg isotope analyses of CV3 chondrules

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Abstract

We have analysed the petrography, major element abundances and bulk Al-Mg isotope systematics of 19 ferromagnesian chondrules from the CV3 chondrites Allende, Mokoia, and Vigarano, together with an Al-rich chondrule and refractory olivine from Mokoia. Co-variations of Al/Mg with Na/Mg and Ti/Mg in our bulk chondrules suggest their compositions are dominantly controlled by reworking of different proportions of chondrule components (e.g. mafic minerals and mesostasis); their precursors are thus fragments from prior generations of chondrules. Our samples show a range in fractionation corrected $^{26}\text{Mg}/^{24}\text{Mg}$ ($\Delta^{26}\text{Mg}$) ~ 60 ppm, relative to precisions $< \pm 5$ ppm (2se) and these values broadly covary with $^{27}\text{Al}/^{24}\text{Mg}$. The data can be used to calculate model initial $^{26}\text{Al}/^{27}\text{Al}$, or $(^{26}\text{Al}/^{27}\text{Al})_0$, of the chondrule precursors. Our resolvably radiogenic chondrules yield model $(^{26}\text{Al}/^{27}\text{Al})_0 \sim 1\text{--}2 \times 10^{-5}$, equivalent to model “ages” of precursor formation ≤ 1 Ma post CAI. However, many of our chondrules show near solar $\Delta^{26}\text{Mg}$ and no variability despite a range in $^{27}\text{Al}/^{24}\text{Mg}$. This suggests their derivation either from younger precursor chondrules or open system behaviour once ^{26}Al was effectively extinct ($(^{26}\text{Al}/^{27}\text{Al})_0 < 0.8 \times 10^{-5}$, given the resolution here). Evidence for the latter explanation is provided by marked rims of orthopyroxene replacing olivine, indicating reaction of chondrules with a surrounding silicate vapour. Concurrent isotopic exchange of Mg with a near chondritic vapour during late reworking could explain their isotopic systematics. One ferromagnesian object is dominated by a high Mg# olivine with elevated Ti and Ca abundances. This refractory olivine has a markedly negative $\Delta^{26}\text{Mg} = -16 \pm 3$ ppm (2se), reflecting its early removal (model age of < 0.5 Ma post CAI), from a reservoir with evolving $\Delta^{26}\text{Mg}$. If representative of the chondrule forming region, this grain defines a minimum interval of radiogenic ingrowth for CV chondrites commensurate with $(^{26}\text{Al}/^{27}\text{Al})_0 > 3.4 \pm 0.6 \times 10^{-5}$. Overall, our samples record a sequence of events from the formation of ferromagnesian objects within 0.5Ma of CAI to re-equilibration of chondrules and

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